How to sit Maxwell and Higgs on the boundary of AdS





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ORGANIZERS

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Based on

arXiv:2211.01760 [pdf, other] hep-th doi 10.1007/JHEP02(2023)012

Holography and magnetohydrodynamics with dynamical gauge fields

Authors: Yongjun Ahn, Matteo Baggioli, Kyoung-Bum Huh, Hyun-Sik Jeong, Keun-Young Kim, Ya-Wen Sun

Abstract: Within the framework of holography, the Einstein-Maxwell action with Dirichlet boundary conditions corresponds to a dual conformal field theory in presence of an external gauge field. Nevertheless, in many real-world applications, e.g., magnetohydrodynamics, plasma physics, superconductors, etc. dynamical gauge fields and Coulomb interactions are fundamental. In this work, we consider bottom-up ho... ⊽ More
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Comments: 54 pages, 22 figures
Journal ref: JHEP02(2023)012

arXiv:2302.02364 [pdf, other] hep-th

Collective dynamics and the Anderson-Higgs mechanism in a bona fide holographic superconductor

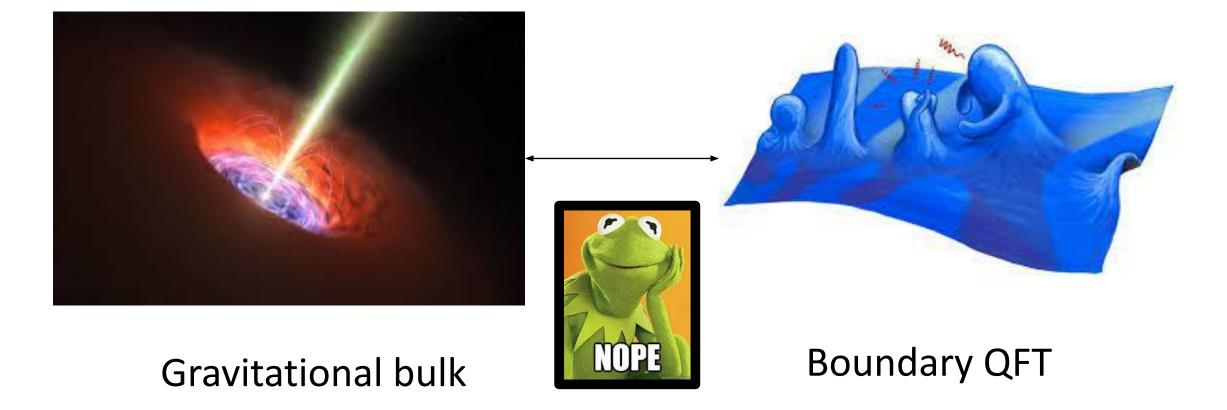
Authors: Hyun-Sik Jeong, Matteo Baggioli, Keun-Young Kim, Ya-Wen Sun

Abstract: The holographic superconductor is one of the most popular models in the context of applied holography. Despite what its name suggests, it does not describe a superconductor. On the contrary, the low temperature phase of its dual field theory is a superfluid with a spontaneously broken U(1) global symmetry. As already observed in the previous literature, a bona fide holographic superconductor can b... ∇ More

Submitted 5 February, 2023; originally announced February 2023.

Comments: 43 pages, 18 figures

Holography



HolographyGravitational bulk action
$$S = \int d^{d+1}x \sqrt{-g} \left[R - 2\Lambda + \mathcal{L}(F_{\mu\nu}) + \dots\right]$$
Boundary
conditions Λ $\Lambda_{\mu}(r,t,\vec{x}) \underset{r \to \infty}{\sim} A_{\mu}^{(0)}(t,\vec{x}) + A_{\mu}^{(1)}(t,\vec{x}) r^{1-D}$

A typical example

- Temperature
- Charge
- Magnetic field

$$BULK: \qquad S = \int d^4x \sqrt{-g} \left(R + 6 - \frac{1}{4}F^2\right), \qquad F = dA,$$

$$ds^2 = -f(r) dt^2 + \frac{1}{f(r)} dr^2 + r^2(dx^2 + dy^2), \qquad A = A_t(r) dt - \frac{B}{2}y dx + \frac{B}{2}x dy,$$

$$A_\mu(r, t, \vec{x}) \underset{r \to \infty}{\sim} A_\mu^{(0)}(t, \vec{x}) + A_\mu^{(1)}(t, \vec{x}) r^{1-D}$$

$$A_{ssuming standard quantization}$$

$$Dirichlet boundary conditions$$

$$Z[g_{\mu\nu}, A_\mu] = \int \mathcal{D}\Phi \exp\left[iS_0(\Phi) + i\int d^3x \left(A_\mu J^\mu(\Phi) + \frac{1}{2}g_{\mu\nu}T^{\mu\nu}(\Phi)\right)\right]$$

What do we get on the other side?

A QFT with a conserved U(1) current in presence of an external gauge field

$$\partial_{\mu}T^{\mu\nu} = F^{\nu\mu}J_{\mu}$$
$$\partial_{\mu}J^{\mu} = 0.$$

Diffusion of conserved charge

Hydrodynamics of the gauge sector $\omega = -iDk^2$

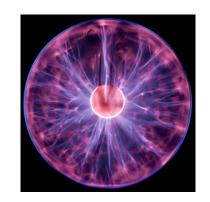


Is this everything we can do?

Unsatisfactory in many physical situations

DYNAMICAL GAUGE FIELDS COULOMB INTERACTIONS









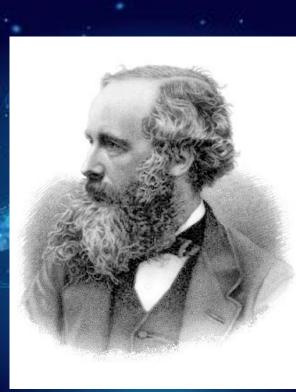
electromagnetism

plasmas

plasmons

superconductors

PART I







Fancy Maxwell's equations

Generalized global symmetries and dissipative magnetohydrodynamics

Sašo Grozdanov, Diego M. Hofman, Nabil Iqbal

Electromagnetism without gauge symmetries

The true global symmetry of U(1) electrodynamics is actually something different. Consider the following antisymmetric tensor

$$J^{\mu\nu} = \frac{1}{2} \varepsilon^{\mu\nu\rho\sigma} F_{\rho\sigma}.$$
 (1.2)

It is immediately clear from the Bianchi identity (i.e. the absence of magnetic monopoles) that $\nabla_{\mu}J^{\mu\nu} = 0$. This is not related to the conservation of electric charge, but rather states that magnetic field lines cannot end.

$$\nabla_{\mu}J^{\mu\nu} = 0.$$

$$\nabla_{\mu}T^{\mu\nu} = H^{\nu}_{\ \rho\sigma}J^{\rho\sigma},$$

Only global symmetries Better for hydrodynamics, etc...

Magnetohydrodynamics from holography

Generalised global symmetries in holography: magnetohydrodynamic waves in a strongly interacting plasma

Sašo Grozdanov, Napat Poovuttikul

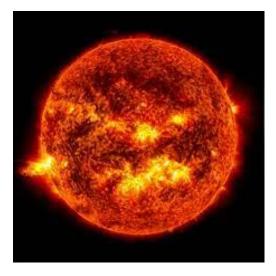
$$\begin{split} S &= \frac{N_c^2}{8\pi^2} \Bigg[\int d^5 x \sqrt{-G} \, \left(R + \frac{12}{L^2} - \frac{1}{3e_H^2} H_{abc} H^{abc} \right) \\ H &= dB, \quad B \to B + d\lambda \\ B_{\mu\nu} \longrightarrow J^{\mu\nu} \end{split} \qquad \begin{array}{l} \text{Short summary:} \\ \text{the standard thing but with} \\ \text{a higher-form} \textcircled{\odot} \end{array} \end{split}$$

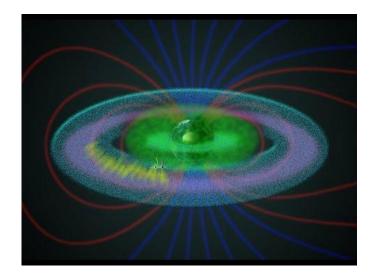
What do they get?

Relativistic magnetohydrodynamics

Juan Hernandez, Pavel Kovtun

A QFT with a dynamical gauge field (dynamical E, B)





Alfven waves

Magnetosonic waves

Our questions/results

- 1. Where is the trick ?
- 2. Do we really need this fanciness ?
- 3. Is there another solution ?
- 4. Magnetohydrodynamics from holography (no dress code)
 - The story of today ...







Yes!

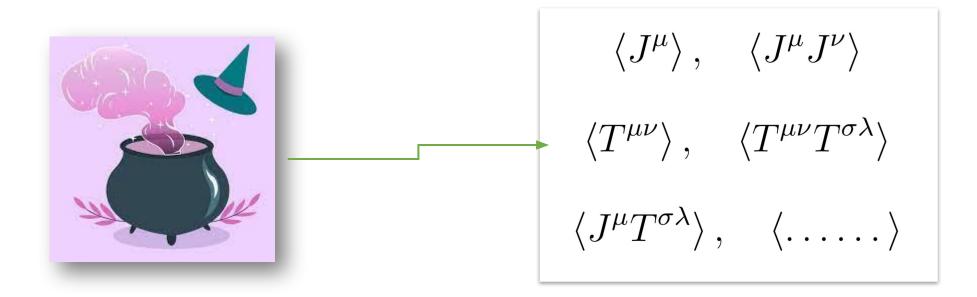
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Field theory break I

Let us start by considering the generating functional $Z[g_{\mu\nu}, A_{\mu}]$:

$$Z[g_{\mu\nu}, A_{\mu}] = \int \mathcal{D}\Phi \exp\left[iS_0\left(\Phi\right) + i\int d^3x \left(A_{\mu}J^{\mu}\left(\Phi\right) + \frac{1}{2}g_{\mu\nu}T^{\mu\nu}\left(\Phi\right)\right)\right],$$

$$S[g_{\mu\nu}, A_{\mu}] := -i \ln Z[g_{\mu\nu}, A_{\mu}]$$



Field theory break II

$$S_{\text{tot}} = S_{\text{m}}[g_{\mu\nu}, A_{\mu}] + \int d^3x \sqrt{-g} \left[-\frac{1}{4\lambda} F^2 + A_{\mu} J_{\text{ext}}^{\mu} \right]$$

Maxwell kinetic term + Legendre transform (coupling to external current)

$$\delta_{A_{\mu}}S_{\text{tot}} = \int d^{3}x \sqrt{-g} \left[J_{\text{m}}^{\mu} - \frac{1}{\lambda} \nabla_{\nu} F^{\mu\nu} + J_{\text{ext}}^{\mu} \right] \delta A_{\mu}.$$

MAXWELL EQUATIONS

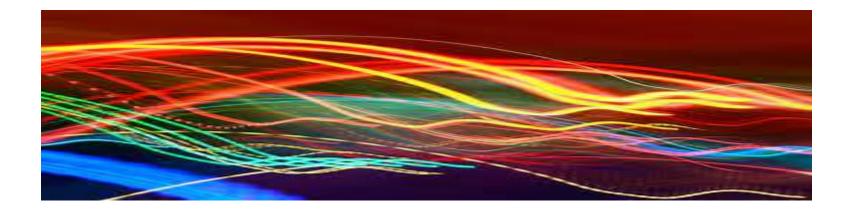
Field theory break III

$$\nabla_{\mu} \left(T_{\rm m}^{\mu\nu} + T_{\rm EM}^{\mu\nu} \right) = F^{\lambda\nu} J_{\rm ext\lambda} , \qquad \qquad \nabla_{\mu} J_{\rm m}^{\mu} = 0 ,$$

$$J_{\rm m}^{\mu} - \frac{1}{\lambda} \nabla_{\nu} F^{\mu\nu} + J_{\rm ext}^{\mu} = 0 , \qquad \qquad \epsilon^{\alpha\beta\gamma} \nabla_{\alpha} F_{\beta\gamma} = 0 ,$$

Standard electromagnetism

$$T^{\mu\nu}_{\mathbf{EM}} = \frac{1}{\lambda} F^{\mu\sigma} F^{\nu}{}_{\sigma} - \frac{1}{4\lambda} F^2 g^{\mu\nu}$$



[more details later]

Old but gold

[Witten, Marolf-Ross, ...]

Boundary asymptotics : $A_{\mu}(r,t,\vec{x}) \sim_{r \to \infty} A^{(0)}_{\mu}(t,\vec{x}) + A^{(1)}_{\mu}(t,\vec{x}) r^{1-D}$,

$$\alpha A^{(0)}_{\mu}(t, \vec{x}) + \beta A^{(1)}_{\mu}(t, \vec{x}) = \text{fixed},$$

MIXED BOUNDARY CONDITIONS

$$\beta = 0 \longrightarrow \mathcal{L}_{CFT} + \int d^d x A^{(0)}_{\mu} J^{\mu}$$
 standard

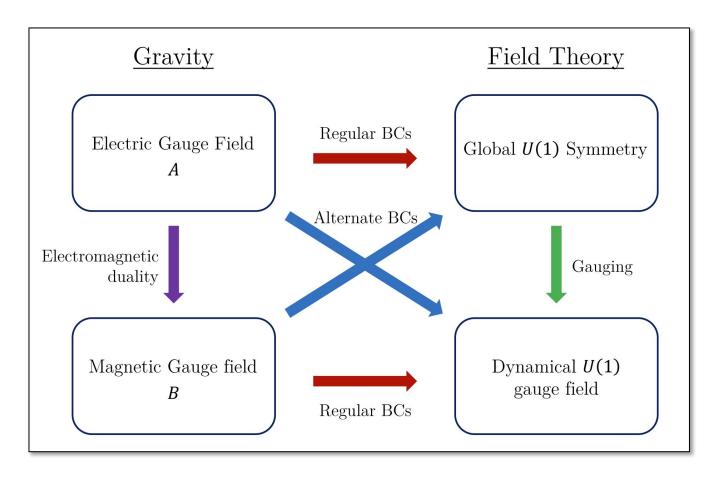
Back to the higher-form trick

arXiv:2010.06594 [pdf, other] hep-th

doi 10.1103/PhysRevD.103.026011

Generalized symmetries and 2-groups via electromagnetic duality in AdS/CFT

Authors: Oliver DeWolfe, Kenneth Higginbotham



Hodge dual in the bulk does not preserve the boundary conditions !!

It changes them from Dirichlet to mixed

Higher-forms are just a fancy way to implement mixed boundary conditions

On the shoulders of giants

$$\Pi^{\mu} - \frac{1}{\lambda} \partial_{\nu} F^{\mu\nu} + J^{\mu}_{\text{ext}} = 0, \qquad \Pi^{\mu} = \frac{\delta S_{\text{on-shell}}}{\delta A_{\mu}} = -\sqrt{-g} F^{r\mu},$$

Our mixed boundary conditions

$$\delta J_{\text{ext}}^{x \ (L)} = Z_{A_1}^{(L)} + \frac{\lambda}{\omega^2 - k^2} Z_{A_1}^{(S)},$$

$$\delta J_{\text{ext}}^{y \ (L)} = (\omega^2 - k^2) Z_{A_2}^{(L)} + \lambda Z_{A_2}^{(S)}.$$

- Lambda parametrizes the EM coupling
- Notice the factors of (w^2-k^2) [back to this later]

Re-discovering the known

$$\epsilon = 2r_h^3 + \frac{\mu^2 r_h}{2} + \frac{B^2}{2r_h} + \frac{B^2}{2\lambda}, \qquad p = r_h^3 + \frac{\mu^2 r_h}{4} - \frac{3B^2}{4r_h} - \frac{B^2}{2\lambda},$$
$$T_{xx} = \left(p + \frac{B^2}{\mu_m}\right) \neq p \qquad T^{\mu}{}_{\mu} = \frac{1}{4\lambda}F^2$$

Thermodynamic and mechanical pressure are not equal

Maxwell theory in 2+1 is scale invariant but not conformal invariant !!

What Maxwell Theory in D<>4 teaches us about scale and conformal invariance

Sheer El-Showk, Yu Nakayama, Slava Rychkov

Magnetohydrodynamics

Step 1: EOMs

$$\begin{vmatrix} \nabla_{\mu} \left(T_{\rm m}^{\mu\nu} + T_{\rm EM}^{\mu\nu} \right) &= F^{\lambda\nu} J_{\rm ext\lambda} , \qquad \nabla_{\mu} J_{\rm m}^{\mu} = 0 , \\ J_{\rm m}^{\mu} - \frac{1}{\lambda} \nabla_{\nu} F^{\mu\nu} + J_{\rm ext}^{\mu} = 0 , \qquad \epsilon^{\alpha\beta\gamma} \nabla_{\alpha} F_{\beta\gamma} = 0 , \end{aligned}$$

Step 2: constitutive relations

$$\begin{split} T^{\mu\nu} &= \epsilon \, u^{\mu} u^{\nu} + p \, \Delta^{\mu\nu} + \mathcal{H}^{\mu\gamma} F^{\nu}_{\ \gamma} + \Pi^{\mu\nu} \,, \\ J^{\mu} &= \rho \, u^{\mu} - \nabla_{\nu} \mathcal{H}^{\mu\nu} + \nu^{\mu} \,, \quad \mathcal{H}^{\mu\nu} := \frac{1}{\lambda} F^{\mu\nu} - M^{\mu\nu}_{\mathbf{m}} \,, \end{split}$$

Step 3: dynamical matrix and QNMs

$$\mathcal{M}(\omega, k) \cdot s_A = 0, \quad \det \mathcal{M}(\omega, k) = 0.$$
$$s_A = \{\delta T, \delta u^{i=x,y}, \delta E^{i=x,y}, \delta B\}$$

Zero density and zero B

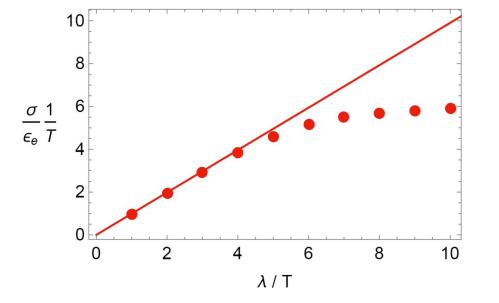
$$\omega \left(\omega + i \frac{\sigma}{\epsilon_{\rm e}} \right) = \frac{k^2}{\epsilon_{\rm e} \,\mu_{\rm m}}$$

$$\omega = -i \frac{\sigma}{\epsilon_{\rm e}} - i \frac{\sigma}{\partial \rho / \partial \mu} k^2$$

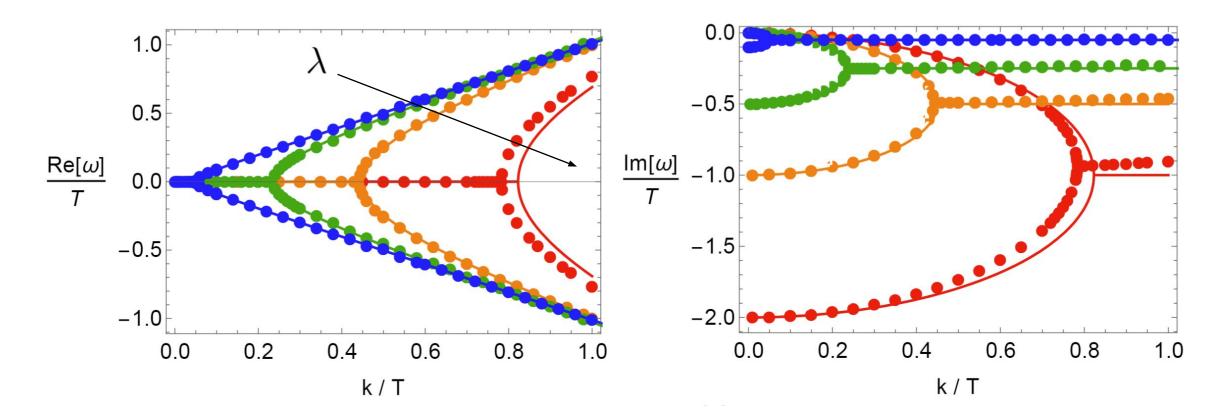
"EM waves"

Damped charge diffusion

We see the effects of polarization and screening (dynamical EM in matter)



Screened EM waves



The photon is screened (skin effect). Just solve it for real w and complex k and back to standard textbooks results

Numerical data (QNMs)
 1st order hydro

Zero density and finite B

All dispersion relations remain of the same type but the coefficients are strongly modified by B^2

$$\omega = \pm v_{\rm ms} \, k - i \frac{\Gamma_{\rm ms}}{2} \, k^2$$

Magnetosonic waves

[sound mode carries magnetic flux now]



Zero density and finite B

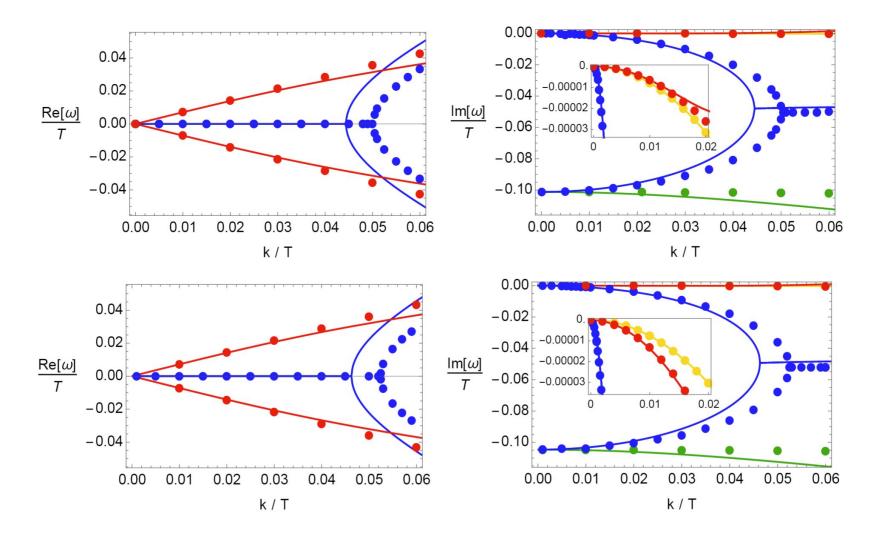


Figure 2. Dispersion relations of the lowest QNMs at zero density $(\mu/T = 0)$ and $B/T^2 \neq 0$. Top and bottom panels are respectively for $B/T^2 = 0.5, 1$.

Finite density and finite B

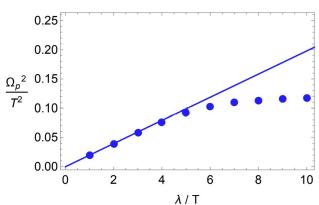
$$\omega = -i \frac{\left(\frac{\partial\rho}{\partial\mu}\right)_{T,B} (\epsilon + p)^2 \sigma}{T \left[\left(\frac{\partial\epsilon}{\partial T}\right)_{\mu,B} \left(\frac{\partial\rho}{\partial\mu}\right)_{T,B} - \left(\frac{\partial\epsilon}{\partial\mu}\right)_{T,B} \left(\frac{\partial\rho}{\partial T}\right)_{\mu,B} \right] (\rho^2 + B^2 \sigma)} k^2, \qquad \omega = -i \frac{\eta}{\mu_{\rm m} \rho^2} k^4.$$

- Longitudinal diffusive mode
- Shear diffusion becomes sub-diffusive
- EM waves and sound modes couple together (4 non-hydro modes)

$$\left[\omega\left(\omega+i\frac{\sigma}{\epsilon_{\rm e}}\right)-\Omega_p^2\right]^2 = \frac{B^2}{\epsilon_{\rm e}^2\,\mu_{\rm m}^2(\epsilon+p)^2}\left[\rho^2-\mu_{\rm m}^2\sigma^2(\rho^2-B^2)+\omega^2\left(2(\epsilon+p)(\sigma-i\epsilon_{\rm e}\omega)\right)\right]\,,$$

where Ω_p is the plasma frequency

$$\Omega_p^2 := \frac{\rho^2}{\epsilon_{\rm e}(\epsilon + p)} \,.$$



Finite density and zero B

$$\left[\omega\left(\omega+i\frac{\sigma}{\epsilon_{\rm e}}\right)-\frac{\Omega_p^2}{\Omega_p^2}\right]^2 = \mathbf{0}$$

$$\left(\sigma^2/\epsilon_{\rm e}^2 \gg 4\,\Omega_p^2\,;\,\,{\rm small \ density}
ight): \quad \omega = -i\,\frac{\epsilon_{\rm e}}{\sigma}\,\Omega_p^2, \qquad \omega = -i\,\frac{\sigma}{\epsilon_{\rm e}} + i\,\frac{\epsilon_{\rm e}}{\sigma}\,\Omega_p^2,$$

 $\left(\sigma^2/\epsilon_{\rm e}^2 \ll 4\,\Omega_p^2\,;\,\,{\rm large \ density}
ight): \quad \omega = \pm\,\Omega_p - i\,\frac{\sigma}{2\epsilon_{\rm e}}\,.$

Small density : overdamped modes Large density : underdamped modes

Finally, setting all the dissipative coefficients (e.g., $\sigma = 0$) to zero, one finds

sound waves
$$\longrightarrow \omega^2 = \Omega_p^2 + v_s^2 k^2$$
, $\omega^2 = \Omega_p^2 + \frac{k^2}{\epsilon_e \mu_m}$. \longleftarrow EM waves

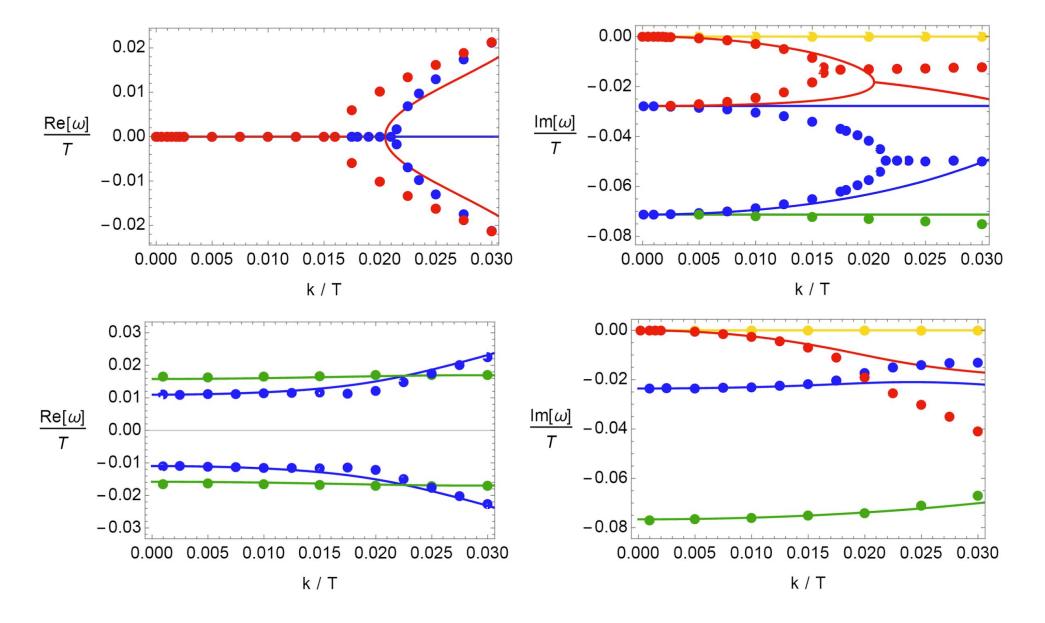


Figure 3. Dispersion relations of the lowest QNMs at finite density ($\mu/T = 0.5$). Top and bottom panels refer respectively to $B/T^2 = 0, 0.5$.

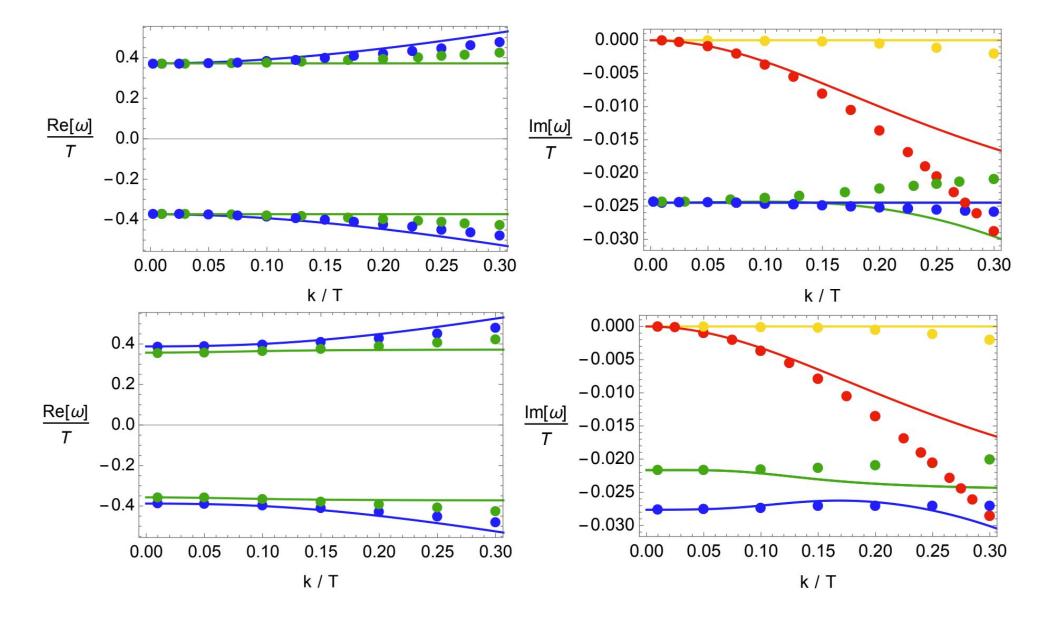


Figure 4. Lowest QNMs at finite density ($\mu/T = 5$). Top and bottom panels refer respectively to $B/T^2 = 0, 0.5$.

Unscreened photon

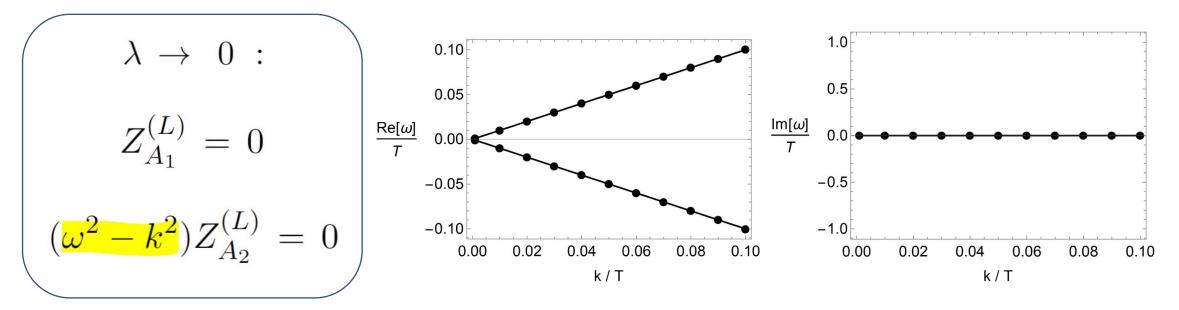


Figure 15. The emergent propagating photon at zero density, zero magnetic field and zero EM coupling $\lambda/T = 0$.



In this limit we are decoupling the photon from matter [no screening, no polarization]: emergent photon

This can be derived analytically [see paper]

Summary of part I

Maxwell + mixed b.c.s.

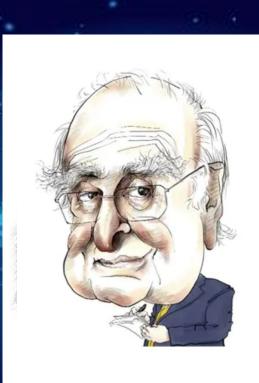


magnetohydrodynamics



[no need of fancy higher forms, sorry]

PART II





Building a Holographic Superconductor

Sean A. Hartnoll, Christopher P. Herzog, and Gary T. Horowitz Phys. Rev. Lett. **101**, 031601 – Published 14 July 2008



Close enough but not!

That and 99% of holographic superconductor papers do not describe a superconductor !!

Making a real holographic superconductor

[Many authors ...]

Holographic Superconductor Vortices

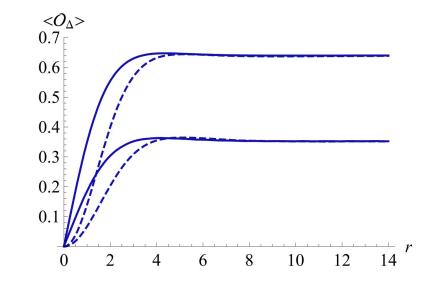
Marc Montull, Alex Pomarol, and Pedro J. Silva Phys. Rev. Lett. **103**, 091601 – Published 26 August 2009

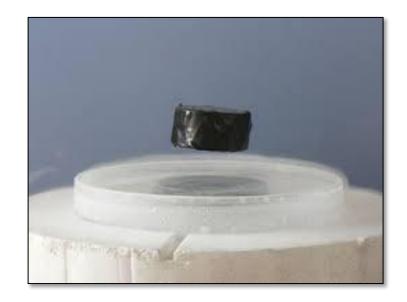
Holographic Meissner Effect

Makoto Natsuume, Takashi Okamura

$$e^2 \mathcal{J}_i = -\frac{1}{\lambda^2} \mathcal{A}_i ,$$

 $\partial_j \mathcal{F}^{ij} = e^2 \mathcal{J}^i .$





SUPERFLUID

1938 Kapitza

Neutral particles

Frictionless flow

Goldstone mode (second sound)

SUPERCONDUCTOR

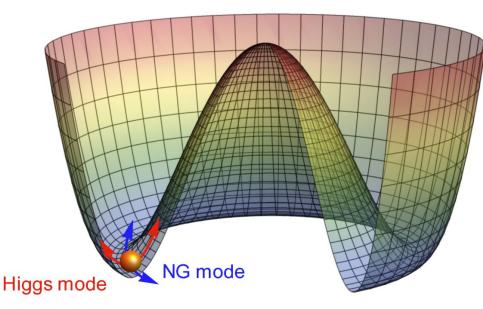
1911 Onnes

Charged particles (electrons)

Flow of current without resistance

Massive gauge field – Meissner effect (Higgs mechanism)

Superfluid Ginzburg-Landau effective theory



Higgs

Complex scalar order parameter

$$\Psi = |\Psi(r)|e^{i\theta(r)}$$

Phenomenological: $b = \beta (T - T_c)$

$$F[\Psi] = F_n(T) + \int d^3r \left[a |\nabla \Psi|^2 + b |\Psi|^2 + \frac{c}{2} |\Psi|^4 \right]$$

mode: $\omega^2 = \frac{2|b|v^2}{a} + v^2 k^2$, Goldstone mode: $\omega^2 = v^2 k^2$

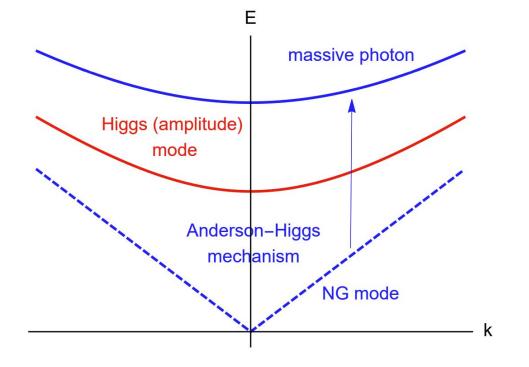
Superfluid Hydrodynamics

Conservation law:
$$\partial_{\mu} J^{\mu} = 0$$
 Superfluid velocity: $u_{\phi} = \frac{1}{\mu} \nabla \phi$
Constitutive relation: $J^{\mu} = \rho_n u^{\mu} + \rho_s u^{\mu}_{\phi} - \sigma_q T \Delta^{\mu\nu} \partial_{\nu} (\mu/T) + \dots$
Josephson relation: $\mu \partial_t u_{\phi} = -\nabla \mu + \xi \rho_s \nabla \nabla \cdot u_{\phi} + \dots$
· Sound mode $v^2 = \frac{\rho_s}{\mu \chi_{\rho\rho}}$
· Infinite conductivity $\sigma = \sigma_q + \frac{i}{\omega} \frac{\rho_s}{\mu}$
· Extended with momentum-energy conservation
· Matching exactly the holographic superfluid model arXiv:2107.08802

Superconductor Ginzburg-Landau effective theory

$$\mathcal{L} = a \left(\partial_{\mu} + i \tilde{q} A_{\mu} \right) \Psi \left(\partial^{\mu} - i \tilde{q} A^{\mu} \right) \Psi^* - b |\Psi|^2 - \frac{c}{2} |\Psi|^4 - \frac{1}{4\lambda} F^2$$

$$|\Psi| = \Psi_0 + \phi, \quad (\phi \ll \Psi_0) \qquad \mathcal{L} = a \,\partial_\mu \phi \,\,\partial^\mu \phi - 2|b| \,\phi^2 + \frac{b^2}{2c} + \left[a \,\tilde{q}^2 \,\Psi_0^2 \,A_\mu A^\mu\right] - \frac{1}{4\lambda} F^2$$



(2) Massive photon:

$$\partial_{\mu}F^{\mu\nu} + \frac{1}{\lambda_{GL}^2}A^{\nu} = 0, \qquad \lambda_{GL}^2 := \frac{v^2}{2a\,q^2\Psi_0^2\,\lambda}$$

Anderson-Higgs mechanism

Holographic superconductor



Same bulk action as for the holographic superfluid model

$$S_{\text{bulk}} = \int \mathrm{d}^4 x \sqrt{-g} \left[R + 6 - \frac{1}{4} F^2 - |D\Psi|^2 - M^2 |\Psi|^2 \right]$$

but different boundary conditions !!

$$\delta J_{\text{ext}}^{x\,(L)} = -\frac{\omega}{\lambda} Z_{A_x}^{(L)} - \frac{1}{e^2} \frac{\omega}{\omega^2 - k^2} Z_{A_x}^{(S)}, \qquad \delta J_{\text{ext}}^{y\,(L)} = -\frac{\omega^2 - k^2}{\lambda} Z_{A_y}^{(L)} - \frac{1}{e^2} Z_{A_y}^{(S)}$$

Already considered in several papers (seminal works: 0906.2396, 1005.1776)

Our focus: collective excitations!



Transverse collective modes

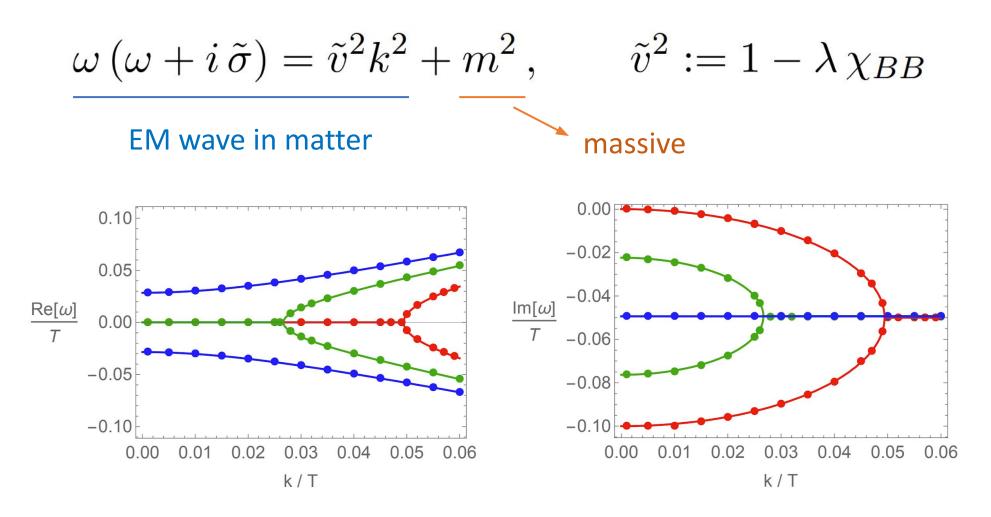
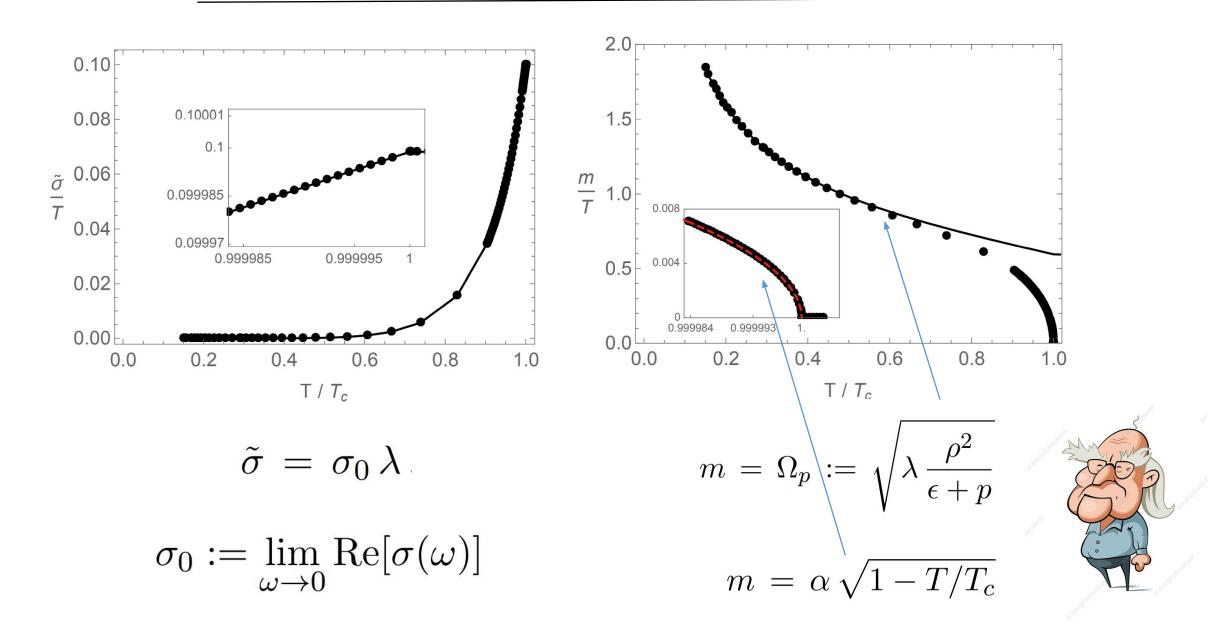
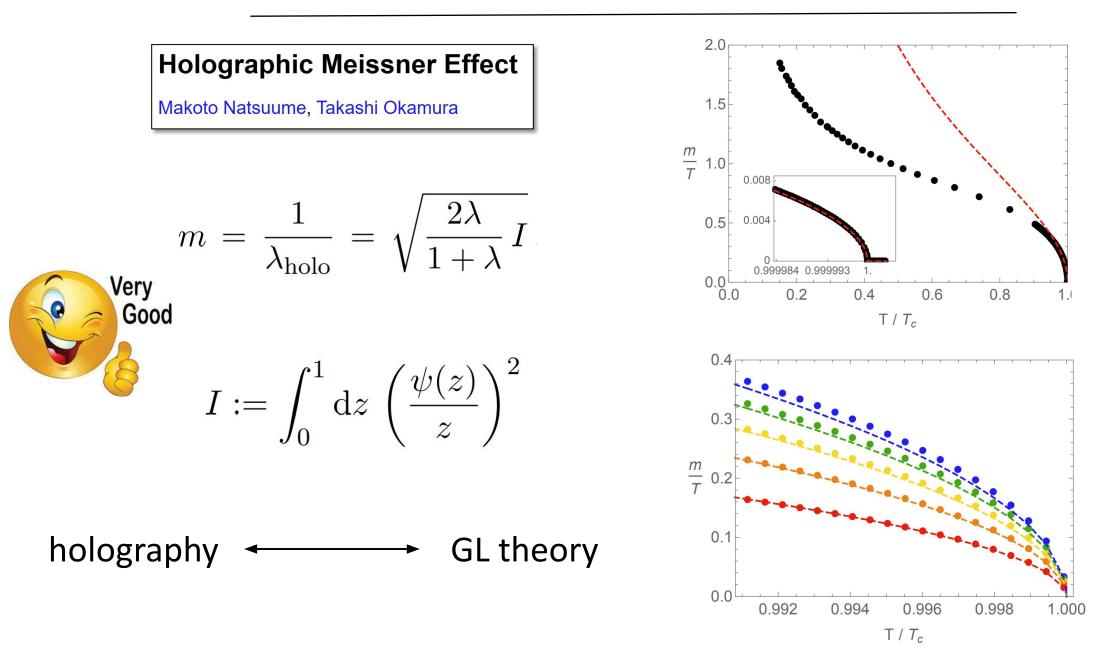


Figure 4. The dispersion relation of the lowest collective modes in the transverse sector for different values of the reduced temperature $T/T_c = (1, 0.999, 0.998)$ (red, green, blue). Symbols represent

Transverse collective modes



Perturbative analytical results



Longitudinal collective modes

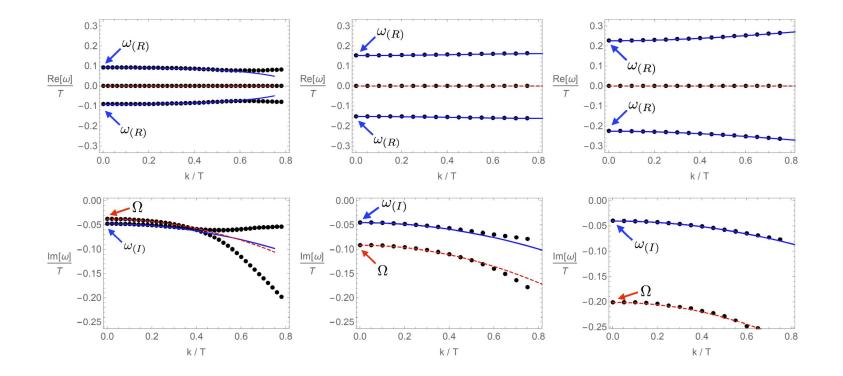
For small wave-vector (very phenomenological)

$$\omega \left(\omega + i \,\tilde{\sigma} + i \,\Gamma \,k^2 \right) = \mathcal{V}^2 \,k^2 + m^2 \,,$$

$$\omega + i\,\Omega + i\,D_\Omega\,k^2 = 0$$

Superfluid sound mode

Superfluid amplitude mode



The fate of the superfluid sound mode

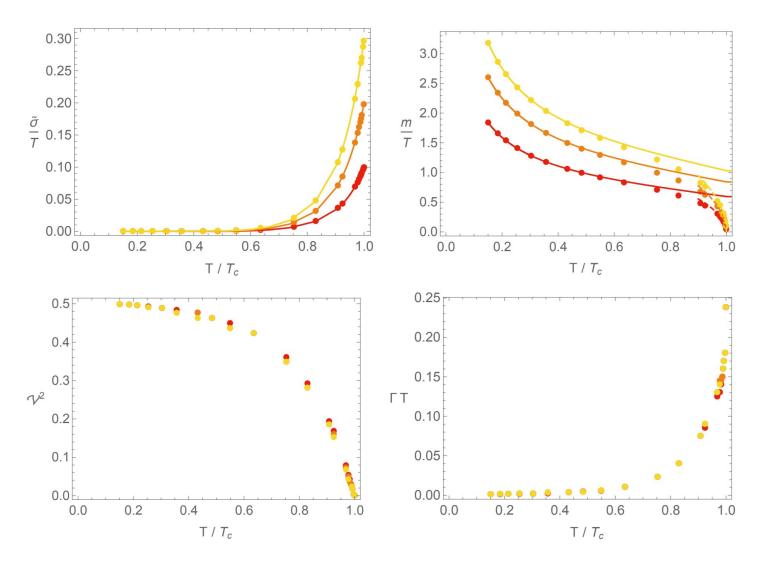


Figure 13. Coefficients appearing in the dispersion relation (6.2) as a function of the reduce temperature for $\lambda = 0.1, 0.2, 0.3$ (red, orange, yellow). Solid lines represent Eq.(5.11) for $\tilde{\sigma}$ and

$$\omega \left(\omega + i \, \tilde{\sigma} + i \, \Gamma \, k^2 \right) = \mathcal{V}^2 \, k^2 + m^2$$

- m is the same as for the transverse mode
- Sigma is the same as for the transverse mode
- Gamma and v do not depend on the Coulomb interaction and they correspond exactly to the speed and the attenuation constant of the superfluid sound mode (0903.2209)

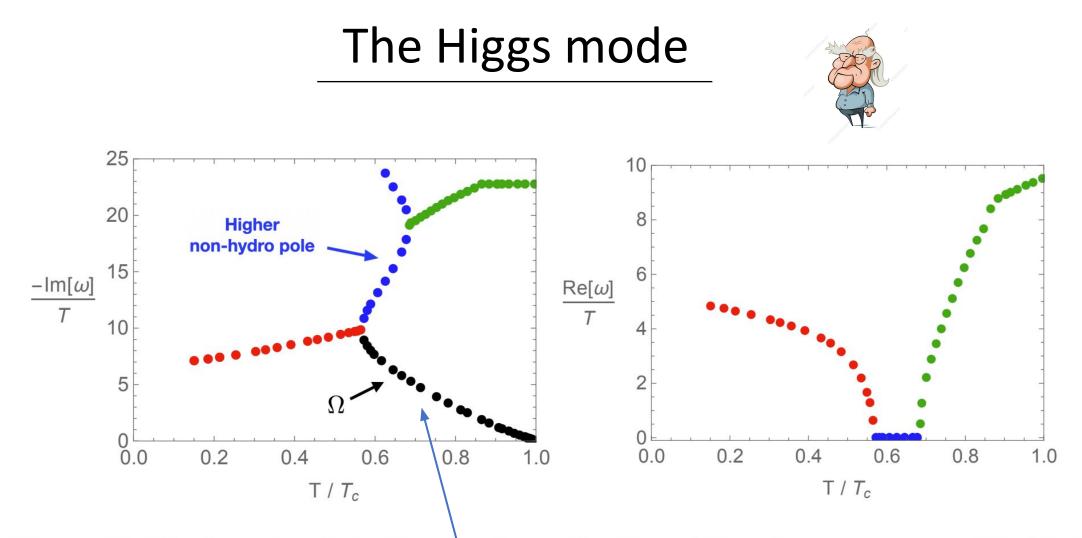
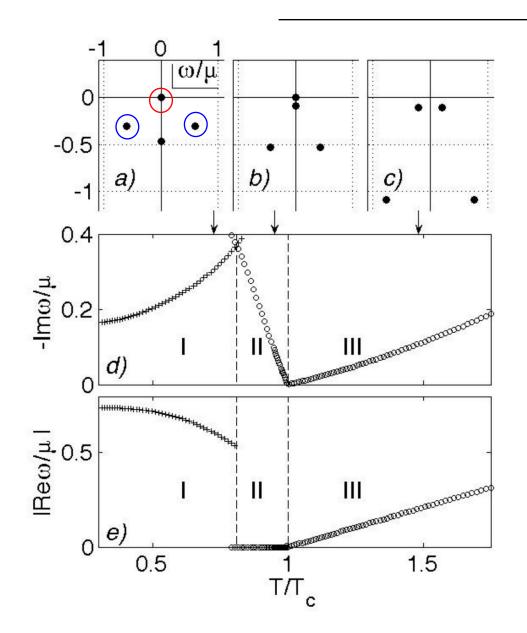


Figure 15. The dynamics of the Higgs mode as a function of the reduce temperature. The black dots encode the fluctuations of the amplitude of the order parameter close to the critical point.

Amplitude mode

VS holographic superfluids



Results from 1207.4194

Overdamped to underdamped crossover but totally different dynamics !

(1) The massive mode at low T is a non-hydro pole which becomes soft:

(2) Always a hydro mode: O

Higgs mode at low temperature

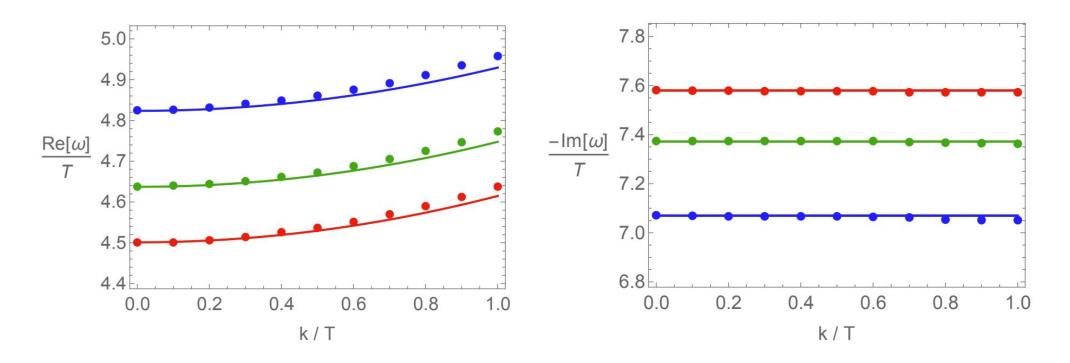
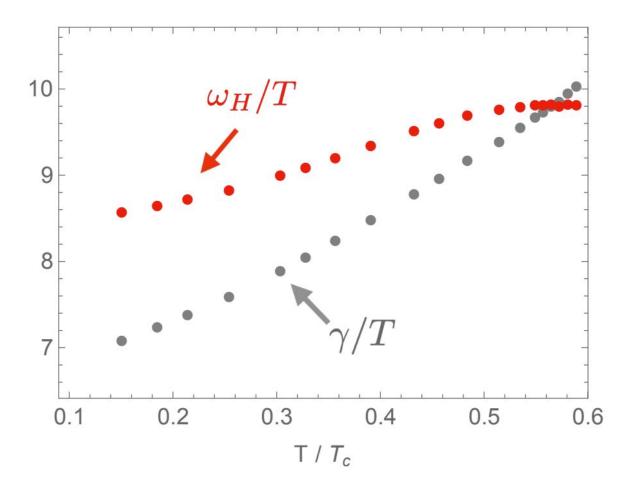


Figure 18. The dispersion relation of the Higgs mode at low temperature: $T/T_c = (0.15, 0.21, 0.25)$ (blue, green, red). Left: Re[ω] vs. k. Right: Im[ω] vs. k. The solid lines are the predictions from GL theory, Eq.(6.9).

Overdamped to underdamped crossover

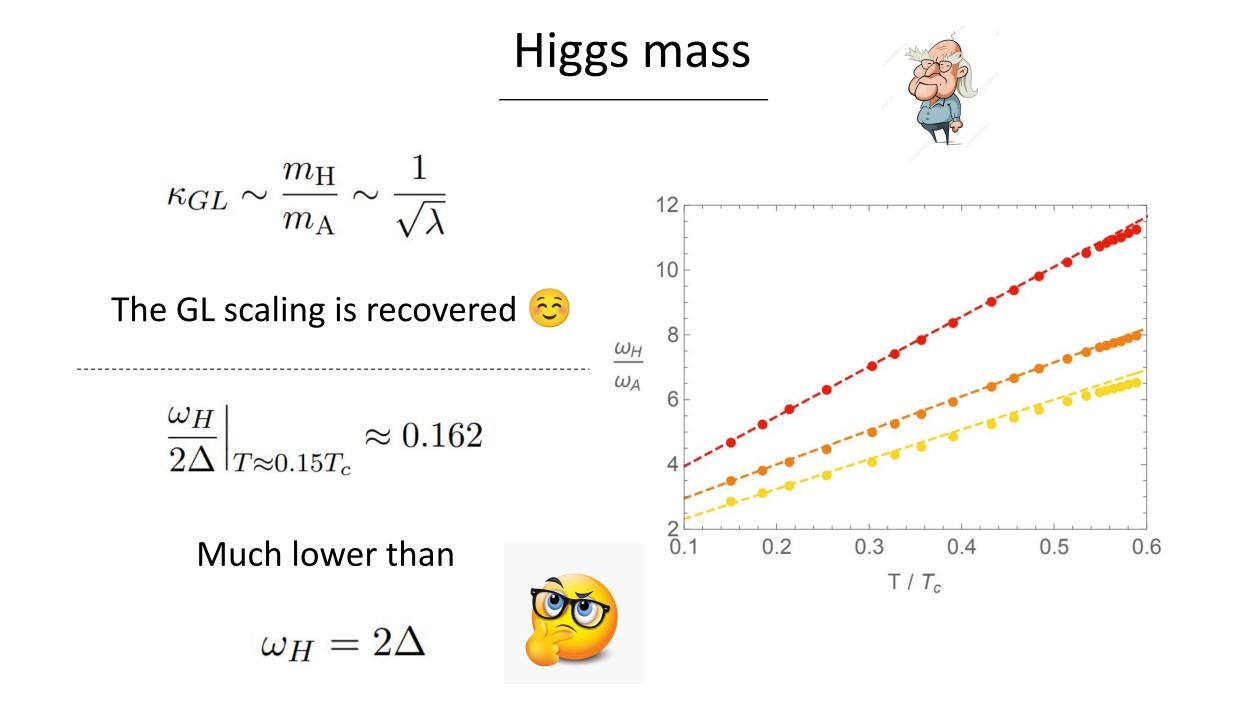
$$\omega^2 = \omega_H^2 + v^2 k^2 - i2\gamma\omega, \qquad \omega_H^2 := \frac{2|b|v^2}{a}$$



$$\omega = -i\gamma \pm \sqrt{\omega_{H}^{2} - \gamma^{2}}$$
 .

Crossover around T=Tc/2

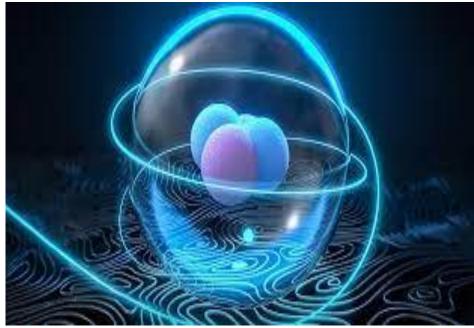
[Interesting aside: dissipative term does seem to vanish at T=0, backreaction ?]



Summary of part II

Hol. superfluid + mixed b.c.s.

Hol. superconductor







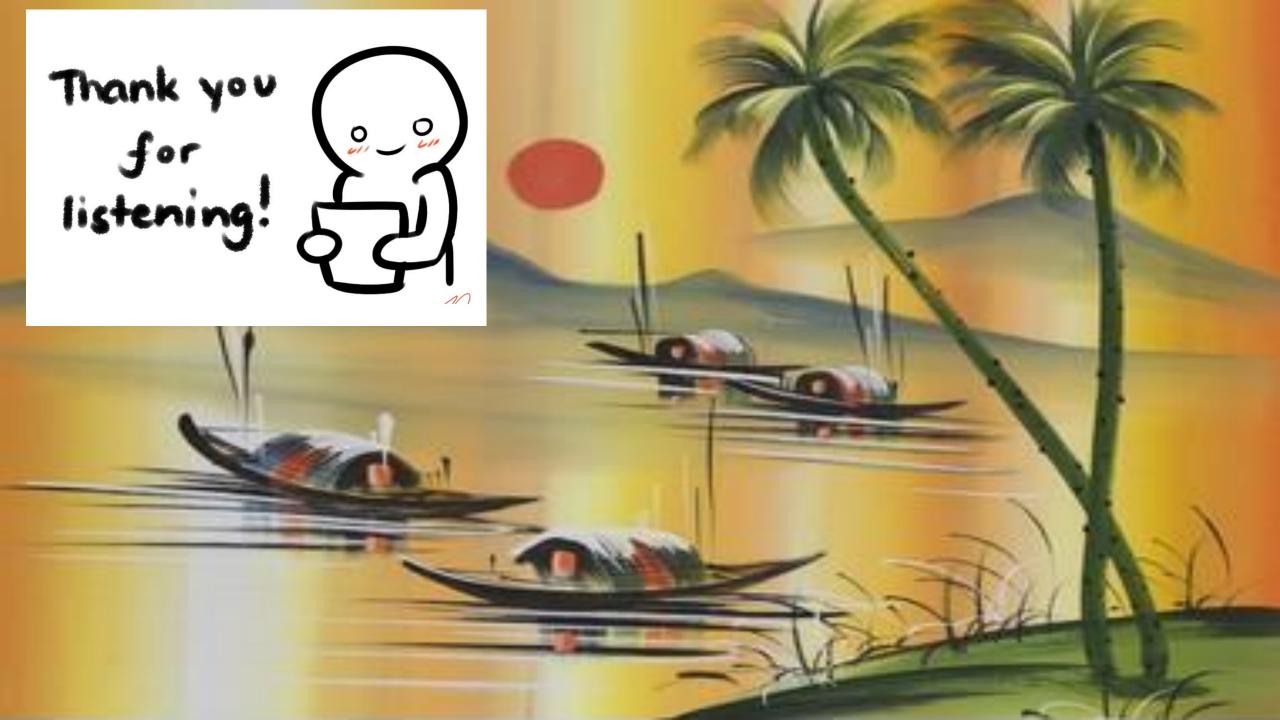


Gauge symmetries in the boundary field theory

Dynamical gauge fields at the boundary

Many applications (cond-mat, plasma physics, QCD, MHD, cosmology)

Relation with bulk higher-form symmetries



more to come...

First application

Building a Holographic Superconductor

Sean A. Hartnoll, Christopher P. Herzog, and Gary T. Horowitz Phys. Rev. Lett. **101**, 031601 – Published 14 July 2008



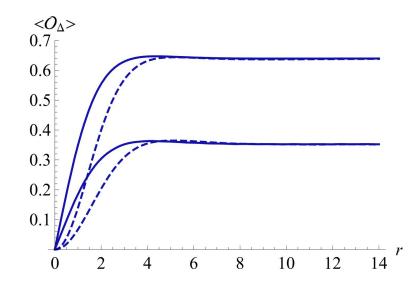
Making a real holographic superconductor (2009)

Example.

[Many authors ...]

Holographic Superconductor Vortices

Marc Montull, Alex Pomarol, and Pedro J. Silva Phys. Rev. Lett. **103**, 091601 – Published 26 August 2009



Recent application

High Energy Physics - Theory

[Submitted on 15 Dec 2017 (v1), last revised 26 Nov 2018 (this version, v4)]

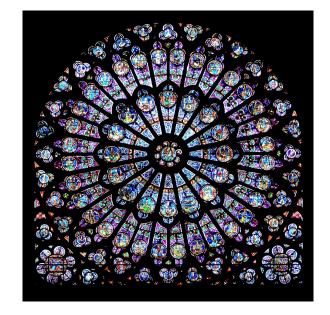
Holographic Plasmons

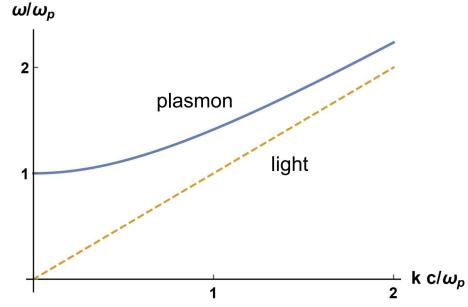
Ulf Gran, Marcus Tornsö, Tobias Zingg

Maxwell equation Continuity equation (Drude model)

$$=\omega_p^2 + c^2 k^2$$

 $\epsilon(\omega) = 1 + rac{4\pi i \sigma(\omega)}{\omega}$





See also

Holographic Meissner Effect

Makoto Natsuume, Takashi Okamura

The holographic superconductor is the holographic dual of superconductivity, but there is no Meissner effect in the standard holographic superconductor. This is because the boundary Maxwell field is added as an external source and is not dynamical. We show the Meissner effect analytically by imposing the semiclassical Maxwell equation on the AdS boundary. Unlike in the Ginzburg-Landau (GL) theory, the extreme Type I limit cannot be reached even in the $e \rightarrow \infty$ limit where e is the U(1) coupling of the boundary Maxwell field. This is due to the bound current which is present even in the pure bulk Maxwell theory. In the bulk 5-dimensional case, the GL parameter and the dual GL theory are obtained analytically for the order parameter of scaling dimension 2.

$$e^2 \mathcal{J}_i = -\frac{1}{\lambda^2} \mathcal{A}_i ,$$

 $\partial_j \mathcal{F}^{ij} = e^2 \mathcal{J}^i .$



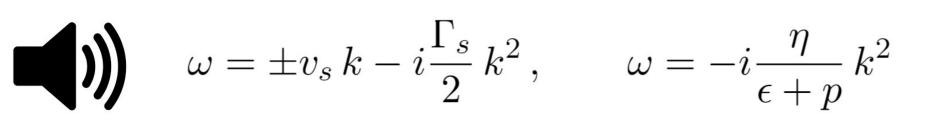
$$\delta_{g_{\mu\nu}}S_{\mathbf{m}} = \frac{1}{2} \int \mathrm{d}^3 x \; \sqrt{-g} \; T_{\mathbf{m}}^{\mu\nu} \, \delta g_{\mu\nu} \,, \quad \delta_{A_{\mu}}S_{\mathbf{m}} = \int \mathrm{d}^3 x \; \sqrt{-g} \; J_{\mathbf{m}}^{\mu} \, \delta A_{\mu} \,.$$

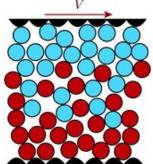
$$\begin{split} T_{\rm m}^{\mu\nu} &= T^{\mu\nu} - T_{\rm EM}^{\mu\nu} = \epsilon_{\rm m} \, u^{\mu} u^{\nu} + p_{\rm m} \, \Delta^{\mu\nu} - M_{\rm m}^{\mu\gamma} \, F^{\nu}_{\,\gamma} + \Pi^{\mu\nu} \\ J_{\rm m}^{\mu} &= J^{\mu} - J_{\rm EM}^{\mu} = \rho_{\rm m} \, u^{\mu} + \nabla_{\nu} M_{\rm m}^{\mu\nu} + \nu^{\mu} \,, \end{split}$$

$$\begin{pmatrix} \frac{\partial p}{\partial B} \end{pmatrix}_{T,\mu} = \left(\frac{\partial p_{\mathbf{m}}}{\partial B} \right)_{T,\mu} - \frac{B}{\lambda} = \chi_{BB}B - \frac{B}{\lambda} ,$$
$$\begin{pmatrix} \frac{\partial \epsilon}{\partial B} \end{pmatrix}_{T,\mu} = \left(\frac{\partial \epsilon_{\mathbf{m}}}{\partial B} \right)_{T,\mu} + \frac{B}{\lambda} = -2\chi_{BB}B + \frac{B}{\lambda}$$

.1 .

Zero density and zero B



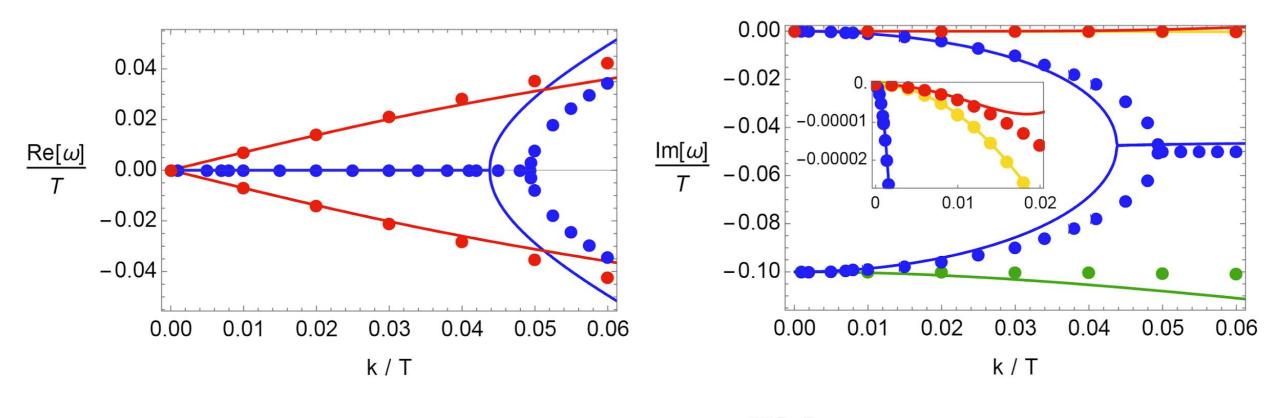


where
$$v_s^2 = \partial p / \partial \epsilon$$
 and $\Gamma_s = \eta / (\epsilon + p)$.

The stress tensor sector is "trivial" since decouples



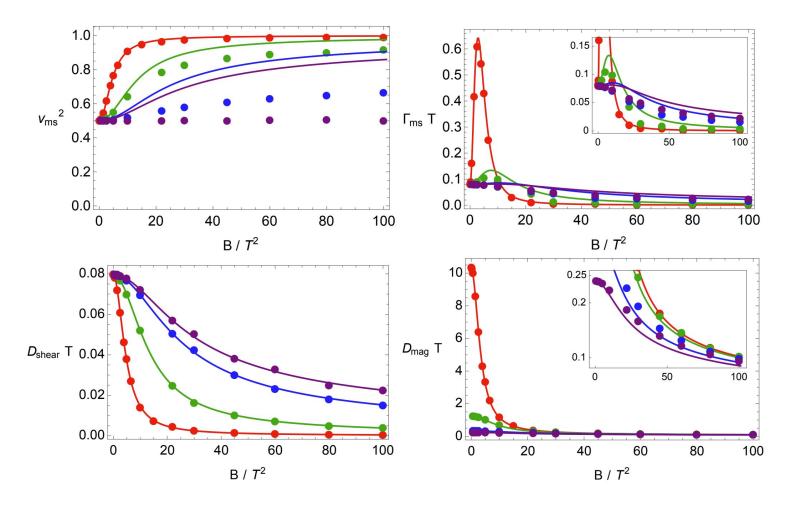
Zero density and zero B



sound waves • EM waves

shear diffusion
 damped charge diffusion

Hydrodynamics breakdown



We see a failure of the hydro description only at large magnetic field together with large EM coupling.

We suspect the problem is the EM coupling



Figure 16. Top: Speed and attenuation constant of magnetosonic waves. Bottom: The diffusion constants of shear and magnetic diffusion. The colors correspond to $\lambda/T = (0.1, 1, 10, 1000)$ (red, green, blue, purple). The insets are a zoom in the low magnetic field regime.

$$\frac{1}{\lambda}\nabla_{\nu}F^{\mu\nu} = J^{\mu}_{\text{free}} + J^{\mu}_{\text{bound}} + J^{\mu}_{\text{ext}}, \qquad (2.18)$$

in which $J^{\mu}_{\text{free}} := \rho_{\text{m}} u^{\mu} + \nu^{\mu}$ and $J^{\mu}_{\text{bound}} := \nabla_{\nu} M^{\mu\nu}_{\text{m}}$. J^{μ}_{free} refers to the current of free charges while J^{μ}_{bound} incorporates the polarization effects. We can decompose the polarization tensor $M^{\mu\nu}$ and $\mathcal{H}^{\mu\nu}$ with respect to fluid velocity as

$$M_{\rm m}^{\mu\nu} = P^{\mu} u^{\nu} - P^{\nu} u^{\mu} - \epsilon^{\mu\nu\rho} u_{\rho} M ,$$

$$\mathcal{H}^{\mu\nu} = u^{\mu} D^{\nu} - u^{\nu} D^{\mu} - \epsilon^{\mu\nu\rho} u_{\rho} H ,$$
(2.19)

and can also be identified with $M_{\rm m}^{\mu\nu} = 2\partial p_{\rm m}/\partial F_{\mu\nu}$, $\mathfrak{H}^{\mu\nu} = 2\partial p/\partial F_{\mu\nu}$. In (3+1) dimensions [50], the magnetization M in Eq.(2.19) becomes the magnetic polarization vector M^{μ} . The electric polarization vector P^{μ} and the magnetization M are associated with the electric field E^{μ} and magnetic field B via the susceptibilities (χ_{EE}, χ_{BB}), i.e.,

$$P^{\mu} = \chi_{EE} E^{\mu}, \qquad M = \chi_{BB} B, \qquad (2.20)$$

with

$$\chi_{EE} = 2 \frac{\partial p_{\rm m}}{\partial E^2}, \qquad \chi_{BB} = 2 \frac{\partial p_{\rm m}}{\partial B^2}.$$
 (2.21)

The physical meaning of D^{μ} and H are the electric displacement vector and the magnetic *H*-field. This can be seen by re-writing Eq. (2.18) in terms of $\mathcal{H}^{\mu\nu}$

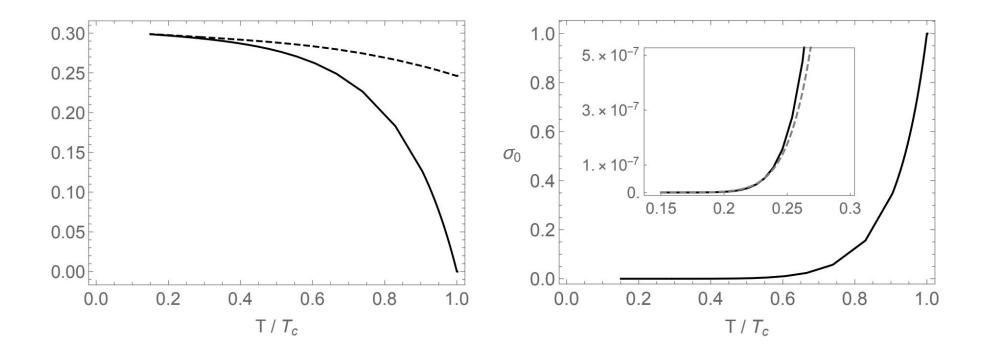
$$\nabla_{\nu} \mathcal{H}^{\mu\nu} = J^{\mu}_{\text{free}} + J^{\mu}_{\text{ext}} \,. \tag{2.22}$$

Eq.(2.20) implies that D^{μ} and H are also proportional to the electric and magnetic field E^{μ} and B via the following relations

$$D^{\mu} = \frac{1}{\lambda} E^{\mu} + P^{\mu} = \epsilon_{e} E^{\mu}, \qquad H = \frac{1}{\lambda} B - M = \frac{1}{\mu_{m}} B, \qquad (2.23)$$

in which we have defined the electric permittivity ϵ_e and the magnetic permeability μ_m . Using all the previous identities and definitions, we finally arrive at the following identities

$$\chi_{EE} = \epsilon_{\rm e} - \frac{1}{\lambda}, \qquad \chi_{BB} = \frac{1}{\lambda} - \frac{1}{\mu_{\rm m}}, \qquad (2.24)$$



the superconducting energy gap Δ via

$$\sigma_0 \sim e^{-\Delta/T}, \qquad \Delta := \sqrt{\langle O_2 \rangle}/2,$$