

# **Hairy AdS black holes with Robin boundary conditions**

Takaaki Ishii (Rikkyo University)

to appear

w/ Tomohiro Harada, Takuya Katagiri, Norihiro Tanahashi

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This can be also considered as

# **Hairy AdS black holes with mixed boundary conditions**

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1. Introduction
2. Global AdS
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# Setup

**Holographic superconductivity:** breaking of U(1) gauge symmetry by Higgs mechanism in AdS

[Hartnoll-Herzog-Horowitz]

"Classic" action ( $8\pi G_N=1$ )

$$S = \int d^4x \sqrt{-g} \left( \frac{1}{2} (R - 2\Lambda) - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - |D\phi|^2 - m^2 |\phi|^2 \right)$$

- Scalar mass:  $m^2 = -2$

Ansatz

$$ds^2 = - (1 + r^2) f(r) e^{-\chi(r)} dt^2 + \frac{dr^2}{(1 + r^2) f(r)} + r^2 d\mathcal{M}_2^2$$

$$A = A_t(r) dt, \quad \phi = \phi(r)$$

- Goal: find solutions with  $\phi \neq 0$ .

# Popular assumptions

We often use **Poincare AdS** and **Dirichlet boundary conditions**.

- Poincare AdS: dual to flat space field theory  $dM_2^2 = dx^2 + dy^2$
- Dirichlet:  $\phi = \frac{\phi_1}{r} + \frac{\phi_2}{r^2} + \dots$  with  $\phi_1 = 0$  in AdS boundary ( $r \rightarrow \infty$ )

## Outcomes

- $\phi \neq 0$  in low temperature, finite charge density
- Spontaneous symmetry breaking  $\langle O_2 \rangle \sim \phi_2 \neq 0$
- Interpretation: superconductivity/superfluidity

# This talk

**Other boundary conditions are available.**

## Often used

- Asymptotically **Poincare AdS**
- **Dirichlet** b.c. for the scalar field  $\phi$

## Options

- Poincare AdS (flat), **global AdS (sphere)**, hyperbolic
- Dirichlet, Neumann, **Robin (mixed b.c.)**

# Preceding work

**Perturbations** of global RNAdS by complex scalar with **Robin b.c.** were studied.

[Katagiri-Harada, 2006.10301]

- Quasinormal modes  $\rightarrow$  (superradiant) instability

**This talk: Backreacted solutions.**

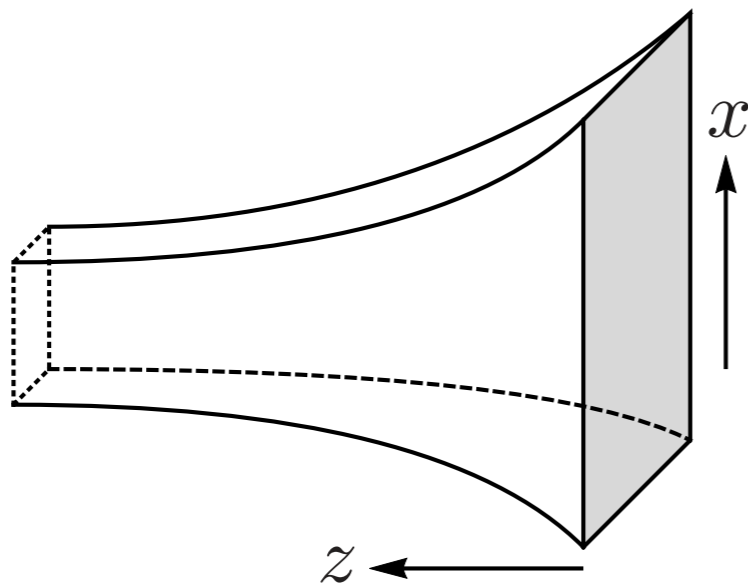


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# Poincare or global AdS

Choices for AdS boundary topology: **Poincare AdS** (flat boundary) or **global AdS** (spherical boundary).



Poincare AdS



Global AdS

**This talk:** We consider asymptotically **global AdS**

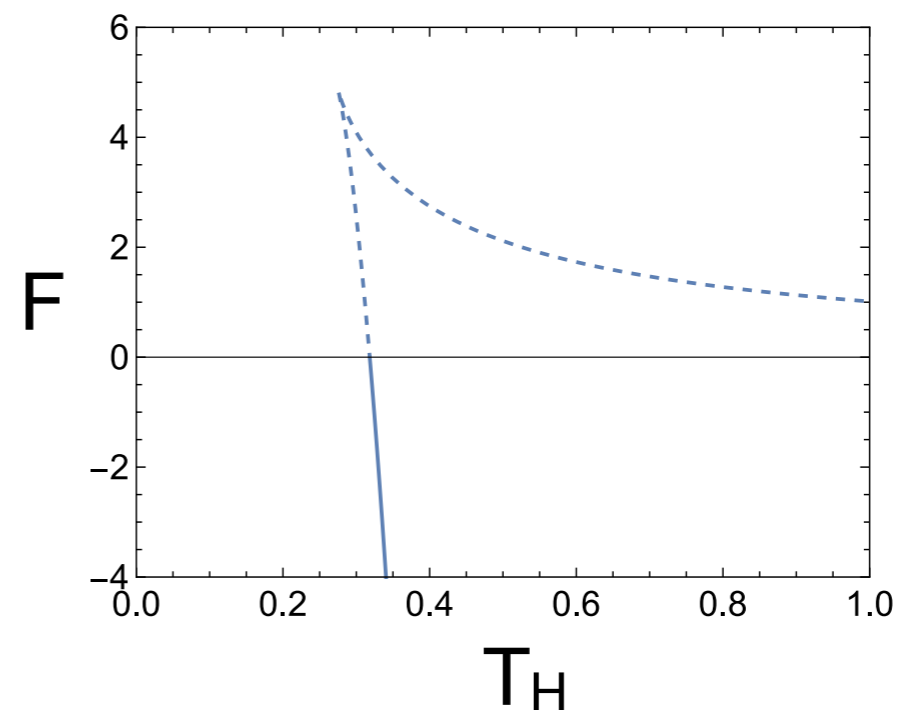
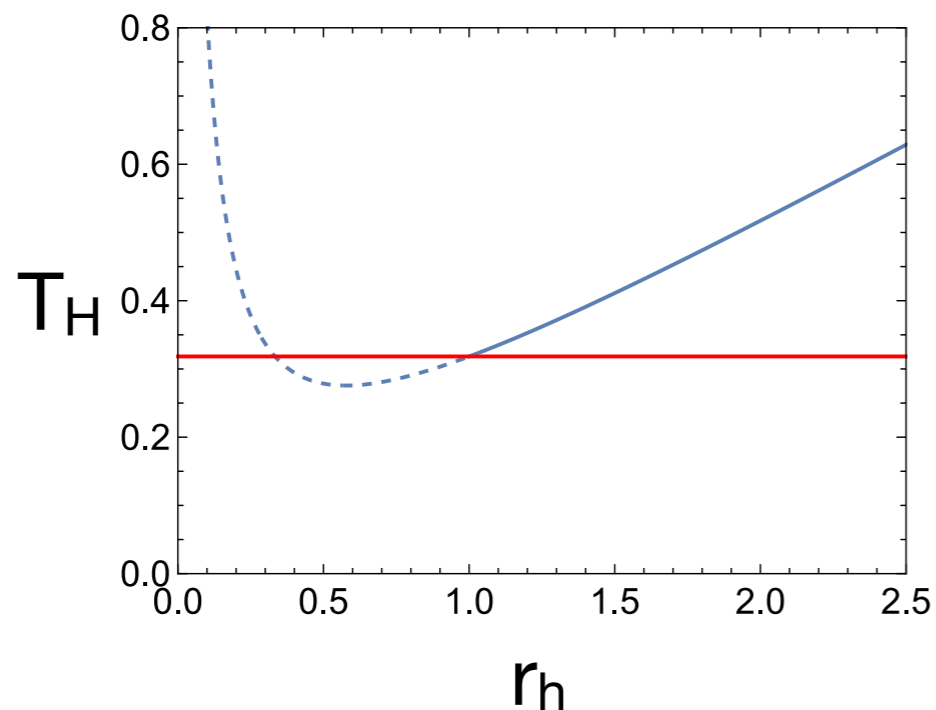
# Hawking-Page phase transition

In global AdS, first order phase transition is discussed between horizonless and BH geometries.

[Hawking-Page, Witten]

- The dominant phase is determined by free energies.

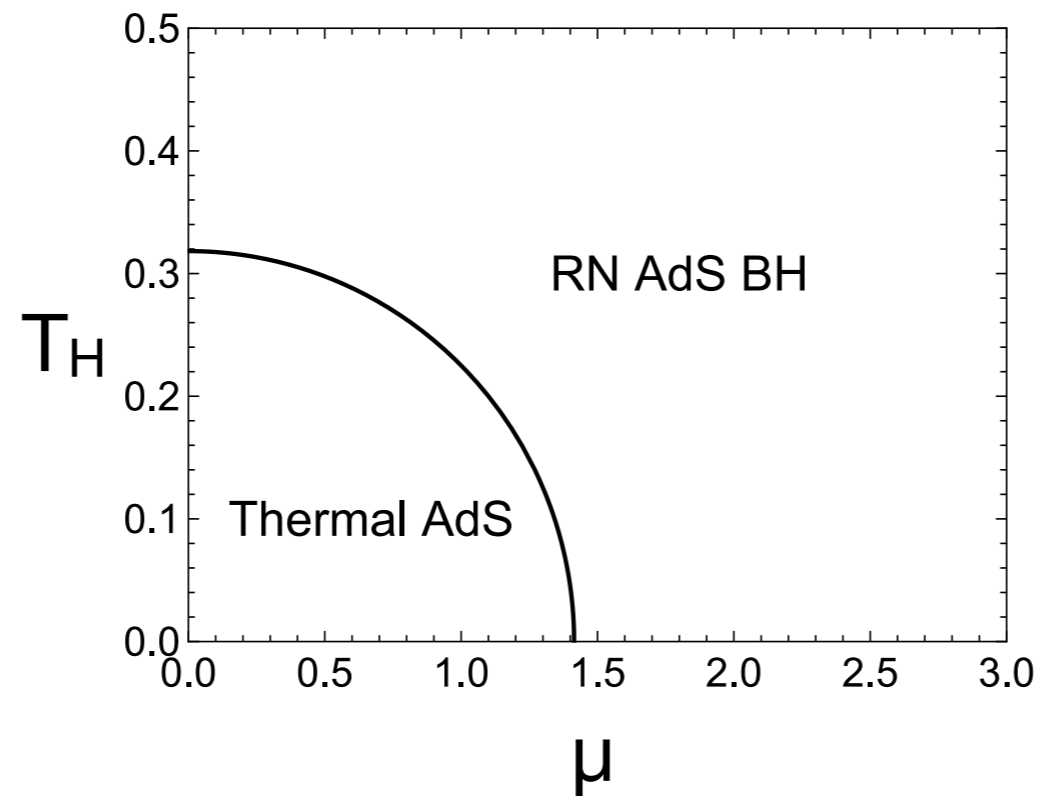
E.g. Thermal AdS vs Schwarzschild AdS BH



# Charged AdS BH

We write phase diagram in **grand canonical ensemble**.  
( $\mu$ : **chemical potential**)

[Chamblin-Emparan-Johnson-Myers]



**This talk:** We also include **charged scalar with Robin boundary conditions**.

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# Scalar boundary conditions

We have several options for the **asymptotic behavior** of the scalar  $\phi$  in  $r \rightarrow \infty$ , if  $-9/4 \leq m^2 \leq -5/4$ .

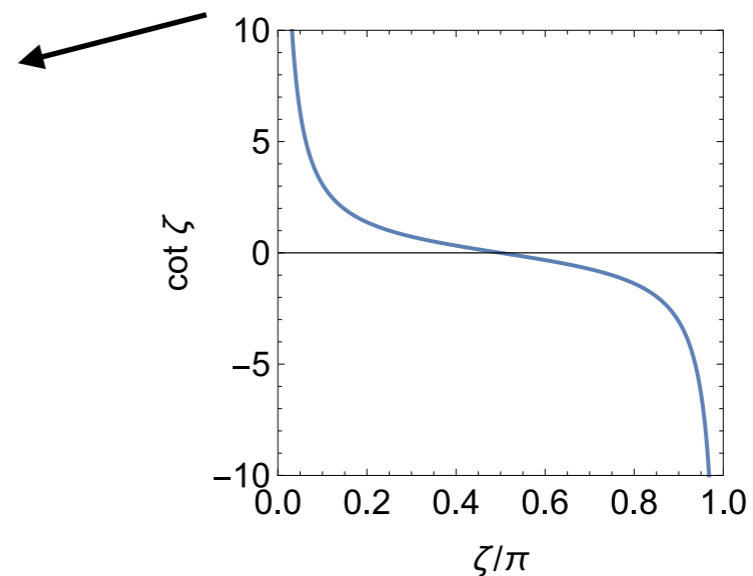
[Witten]

- Asymptotic behavior:  $\phi = \frac{\phi_1}{r} + \frac{\phi_2}{r^2} + \dots$

## Boundary conditions

- Dirichlet  $\phi_1=0$  ( $\zeta=0$ )
- Neumann  $\phi_2=0$  ( $\zeta=\pi/2$ )
- **This talk: Robin  $\phi_1 \neq 0, \phi_2 \neq 0$**

$$\cot \zeta = \frac{\phi_2}{\phi_1} \quad (0 \leq \zeta < \pi)$$



# Hairy solutions ( $\phi \neq 0$ ) in global AdS

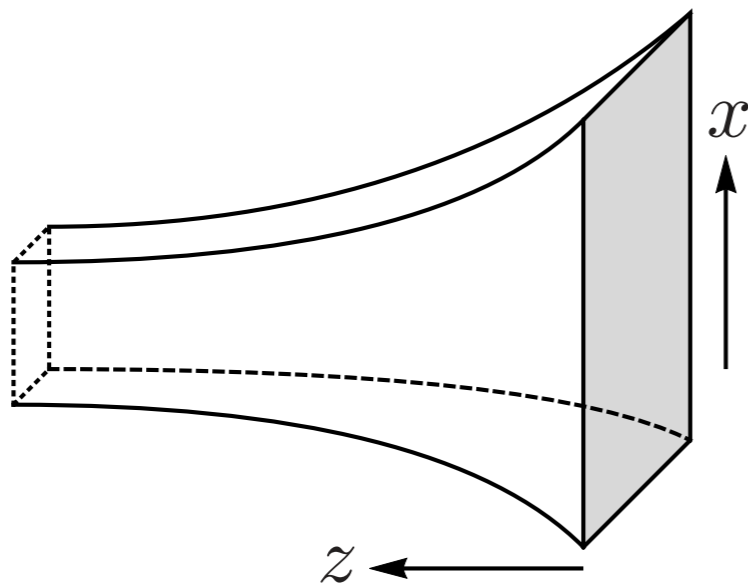
<b>Global AdS</b>	$\mu=0$	$\mu>0$
Horizonless	Neutral boson star	Charged boson star
With BH horizon	Neutral hairy black hole	<b>Charged Hairy black hole</b>

this work

- Parameters:  $(\mu, T_H, \zeta, q)$  ( $q$ : gauge coupling)

# What are boson stars?

In global AdS, **nonzero-scalar horizonless solutions** can be constructed.



Poincare AdS



Global AdS

- Such solutions are called **boson stars**.



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# Holographic renormalization

To obtain the expression of free energy (grand potential), we do **holographic renormalization for Robin b.c.**

[e.g. Skenderis]

- This is a systematic way to obtain such quantities.

3 steps to Robin b.c.

1. Dirichlet: standard addition of counterterms
2. Neumann: Legendre transform from Dirichlet
3. **Robin: modification of Neumann**

# Dirichlet theory

On-shell action diverges, and we make it finite by **adding covariant counterterms.**

[e.g. Skenderis]

## Renormalized action

$$S_{\text{ren}}^D = \lim_{\text{cutoff} \rightarrow \infty} (S_{\text{bulk}} + S_{\text{ct}})$$

## Scalar source and expectation value

$$\Phi_D = \sqrt{2}\phi_1 = 0, \quad \langle O_2 \rangle = \sqrt{2}\phi_2$$

## Variation of action

$$\delta S_{\text{ren}}^D = \int d^3x \sqrt{-h} \left( \frac{1}{2} \langle T^{ij} \rangle \delta h_{ij} + \langle J^i \rangle \delta \Psi_i + \langle O_2 \rangle \delta \Phi_D \right)$$

# Neumann theory

**Legendre transform** of Dirichlet theory by a finite boundary term is Neumann theory.

[Witten, Papadimitriou]

Legendre transform by finite boundary term

$$S_{\text{ren}}^N = S_{\text{ren}}^D + S_{\text{LT}} \quad S_{\text{LT}} = -2 \int d^3x \sqrt{-h} \phi_1 \phi_2$$

Scalar source and expectation value

$$\Phi_N = -\sqrt{2}\phi_2 = 0, \quad \langle O_1 \rangle = \sqrt{2}\phi_1$$

Variation of action

$$\delta S_{\text{ren}}^N = \int d^3x \sqrt{-h} \left( \frac{1}{2} \langle T^{ij} \rangle \delta h_{ij} + \langle J^i \rangle \delta \Psi_i + \langle O_1 \rangle \delta \Phi_N \right)$$

# Double trace deformation

Robin boundary conditions appear as **modification of Neumann theory by a double trace term.**

[Witten, Papadimitriou]

Additional finite term for consistent variation

$$\begin{aligned} S_{\text{ren}}^R &= S_{\text{ren}}^N + S_{\text{Dtr}} \\ &= S_{\text{ren}}^D + S_{\text{LT}} + S_{\text{Dtr}} \end{aligned} \quad S_{\text{Dtr}} = \cot \zeta \int d^3x \sqrt{-h} \phi_1^2$$

Scalar source and expectation value

$$\Phi_R = -\sqrt{2} (\phi_2 - \phi_1 \cot \zeta) = 0, \quad \langle O_1 \rangle = \sqrt{2} \phi_1$$

Variation of action

$$\delta S_{\text{ren}}^R = \int d^3x \sqrt{-h} \left( \frac{1}{2} \langle T^{ij} \rangle \delta h_{ij} + \langle J^i \rangle \delta \Psi_i + \langle O_1 \rangle \delta \Phi_R \right)$$

# Grand potential

We evaluate the phase structure in **grand canonical ensemble** by comparing the grand potential.

Total energy (BH mass)

$$\mathcal{E}_R = \mathcal{E}_N + \mathcal{E}_{\text{Dtr}} = \mathcal{E}_D + \mathcal{E}_{\text{LT}} + \mathcal{E}_{\text{Dtr}}$$

$$\mathcal{E} = \int d\Omega_2 \langle T_{tt} \rangle$$

Grand potential

$$\Omega_R = \Omega_N + \Omega_{\text{Dtr}} = \Omega_D + \Omega_{\text{LT}} + \Omega_{\text{Dtr}}$$

$$\Omega = \mathcal{E} - T_{\text{H}} \mathcal{S}_{\text{BH}} - \mu Q$$

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## 5-1. Neutral case ( $\mu=q=0$ )

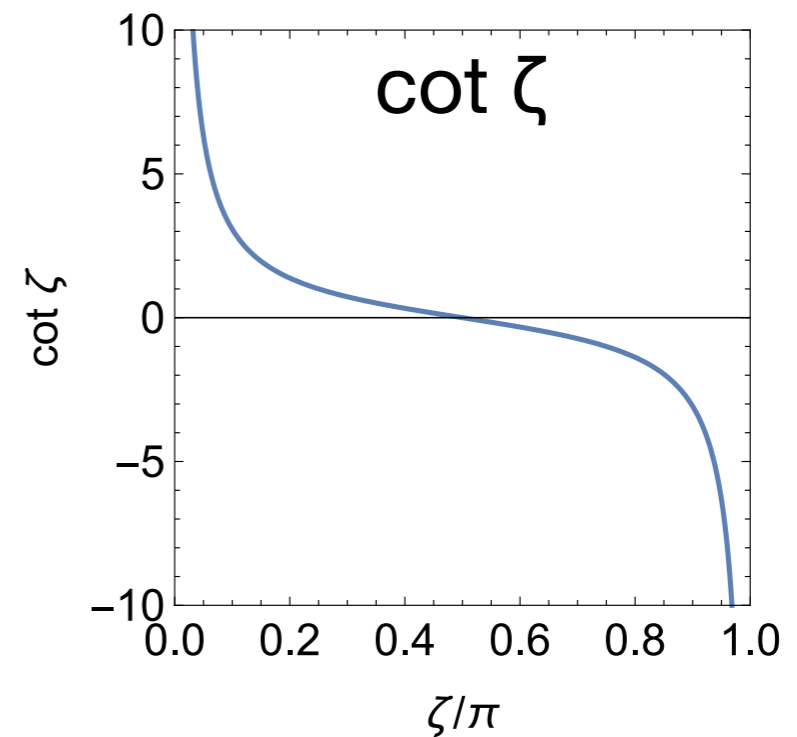
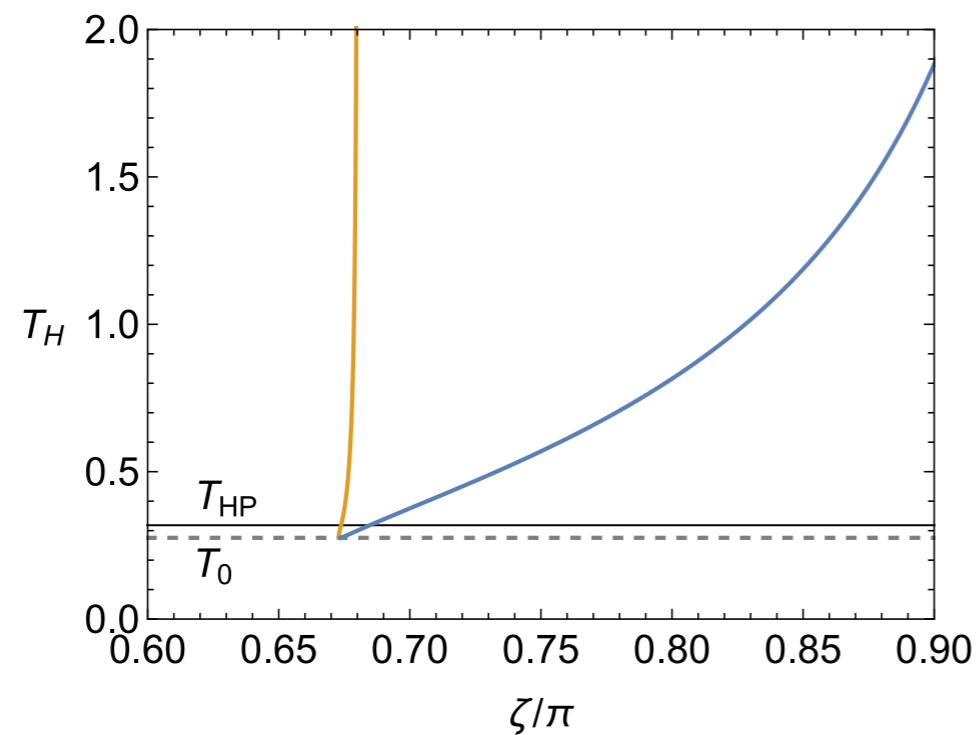
	$\phi=0$	$\phi\neq 0$
No horizon	Thermal AdS	Robin boson star
With BH horizon	Schwarzschild AdS BH	Neutral hairy Robin BH

- We compare free energies of these 4 solutions and determine the phase diagram in the  $(\zeta, T_H)$ -plane.



# Instability for forming scalar hair

The Robin boundary conditions make neutral scalar unstable to become nonzero in  $\zeta/\pi > 0.67$ .

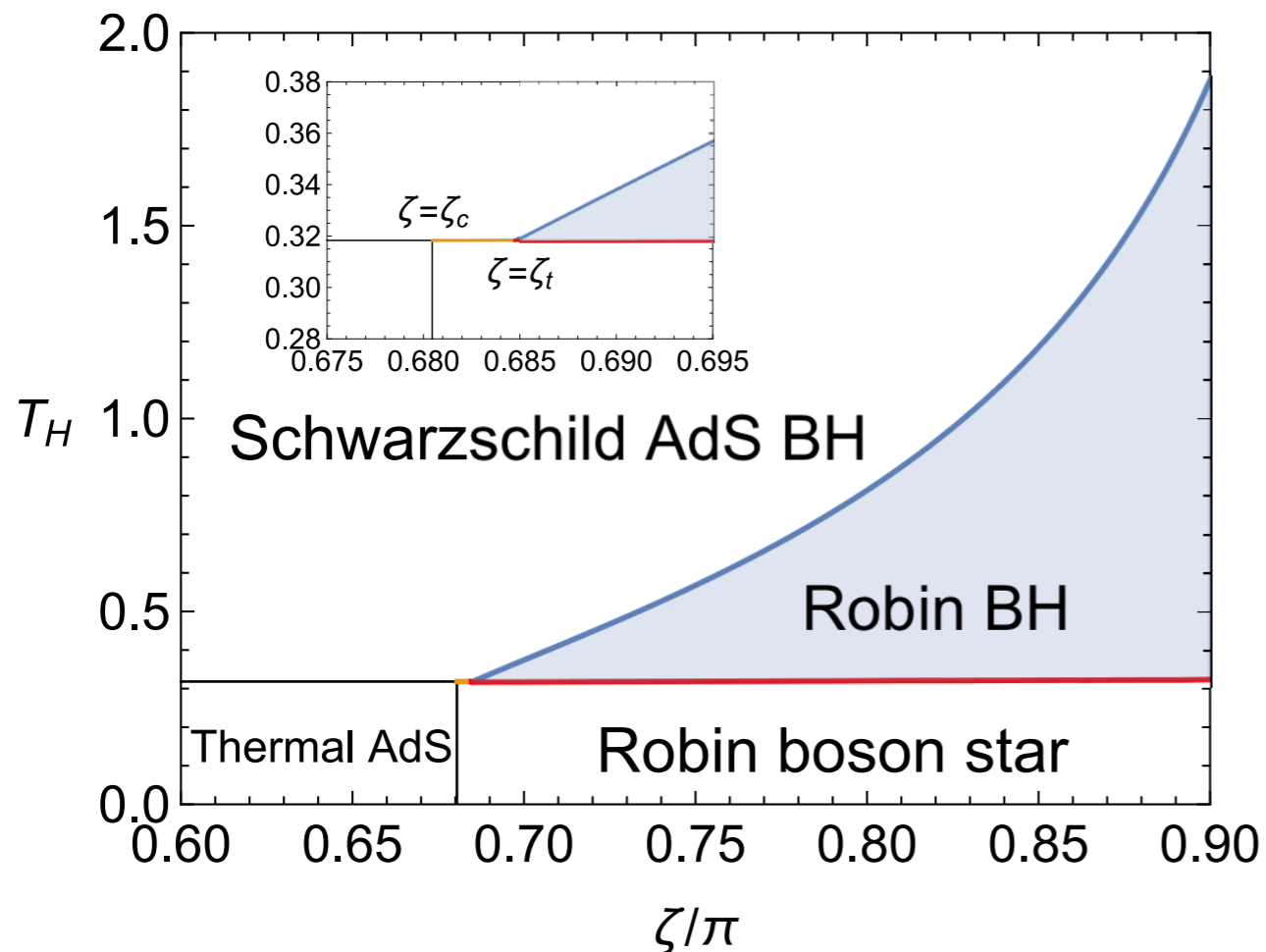


- Figure: location of the onset of instability
- Hairy BH to the right of each curve.

(Orange: small BH branch, Blue: large BH branch)

# Neutral phase diagram

We compare free energies between 4 different solutions: scalar field  $\phi=0$  or  $\phi\neq 0$  (Robin), and BH or horizonless



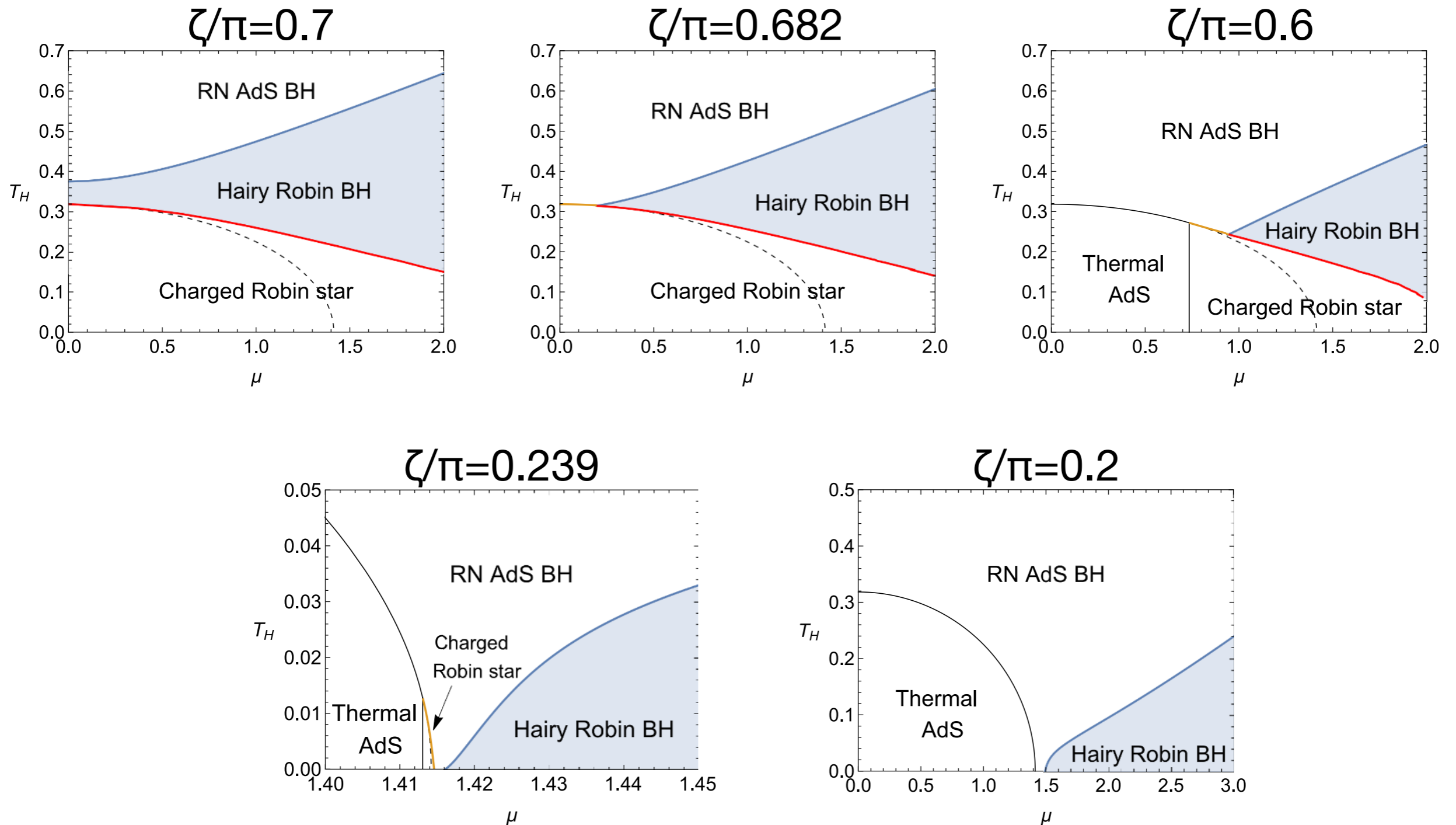
- Blue line: 2nd order, scalar hair
- Red: 1st order Hawking-Page
- Orange: 1st order Hawking-Page

## 5-2. Charged case $(\mu, T_H, \zeta, q) \neq 0$

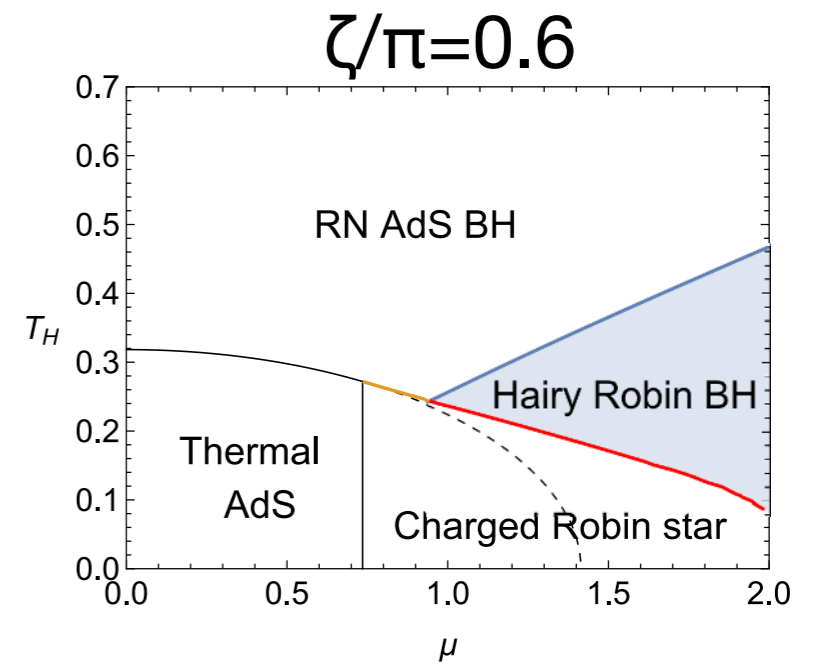
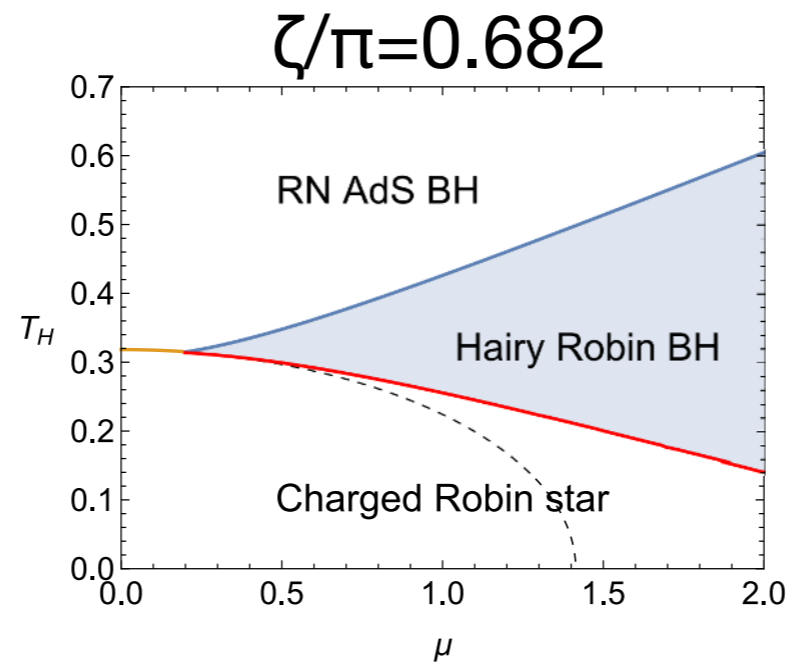
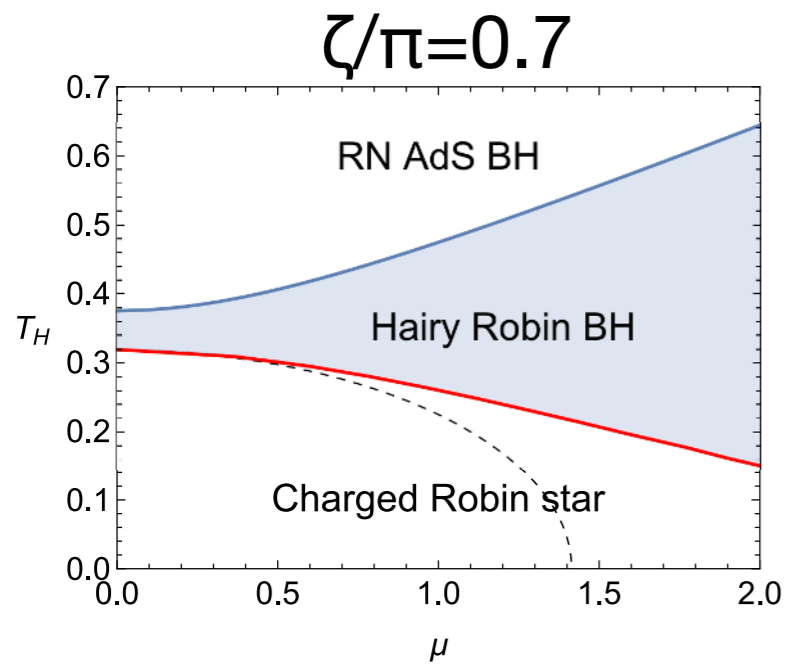
	$\phi=0$	$\phi \neq 0$
No horizon	Thermal AdS	Charged Robin star
With BH horizon	RNAdS BH	Charged Hairy Robin BH

- We compare grand potentials of these 4 solutions and determine the phase diagram in the  $(\mu, T_H)$ -plane.

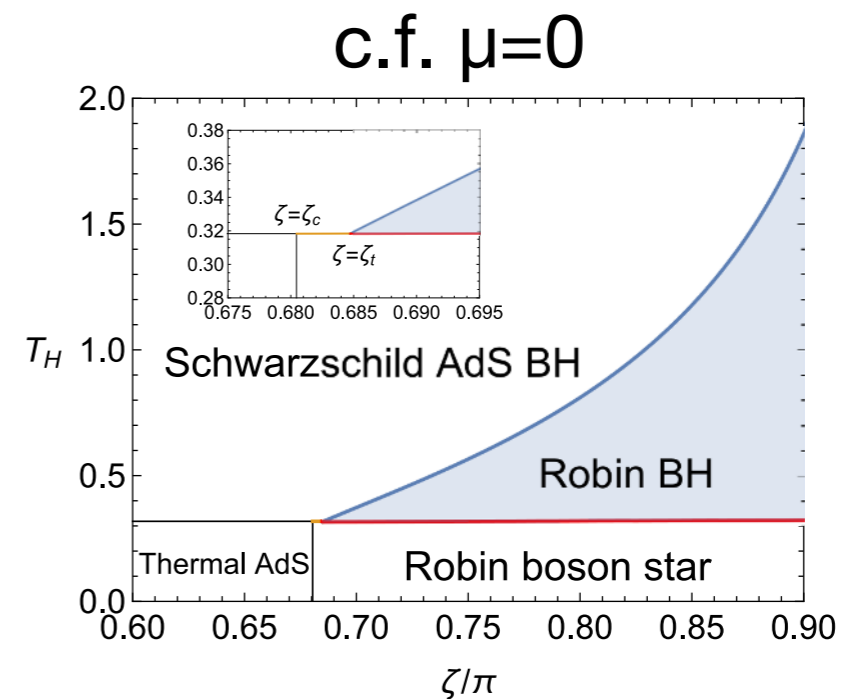
# Charged phase diagrams ( $q=1$ )



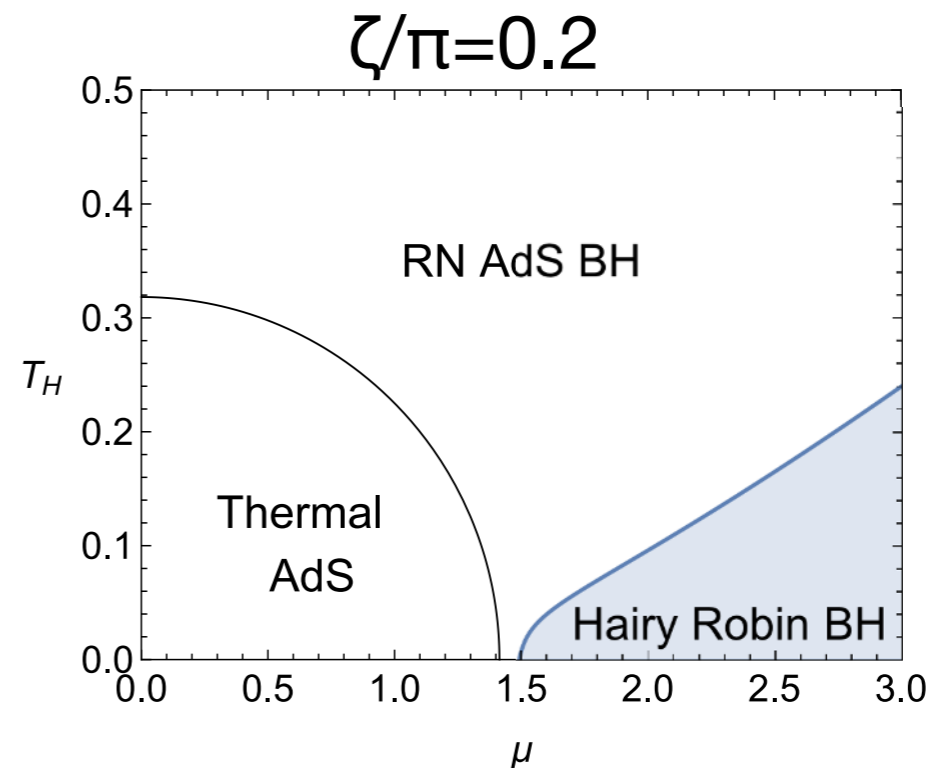
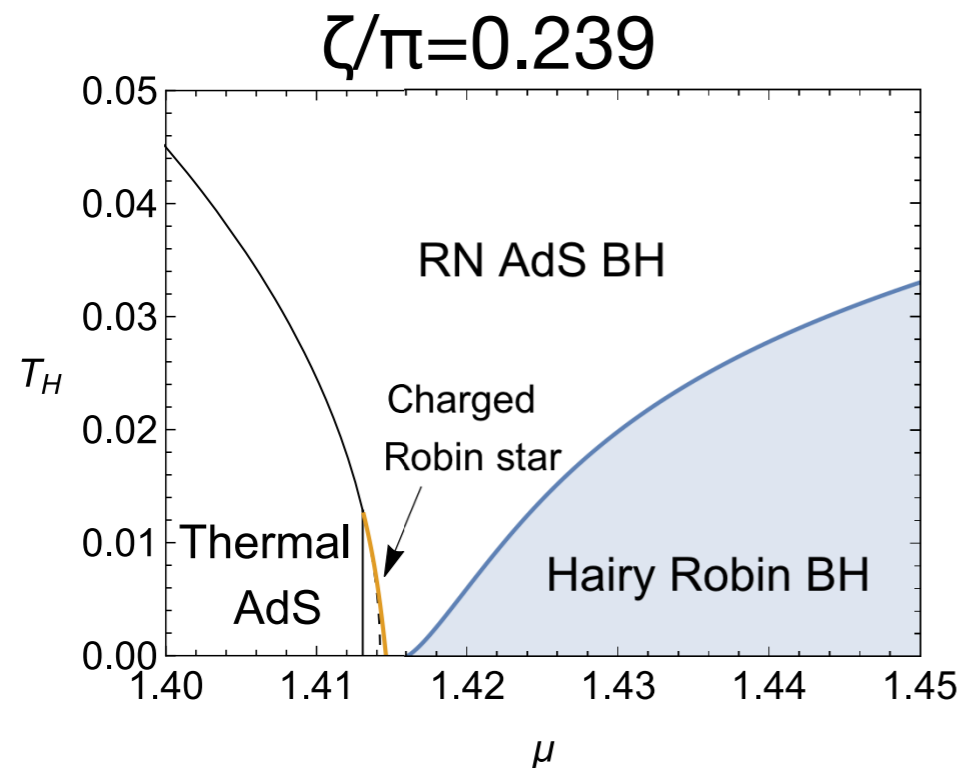
# Large $\zeta$



- Blue line: 2nd order, scalar hair
- Red: 1st order Hawking-Page
- Orange: 1st order Hawking-Page



# Small $\zeta$ at small coupling $q$

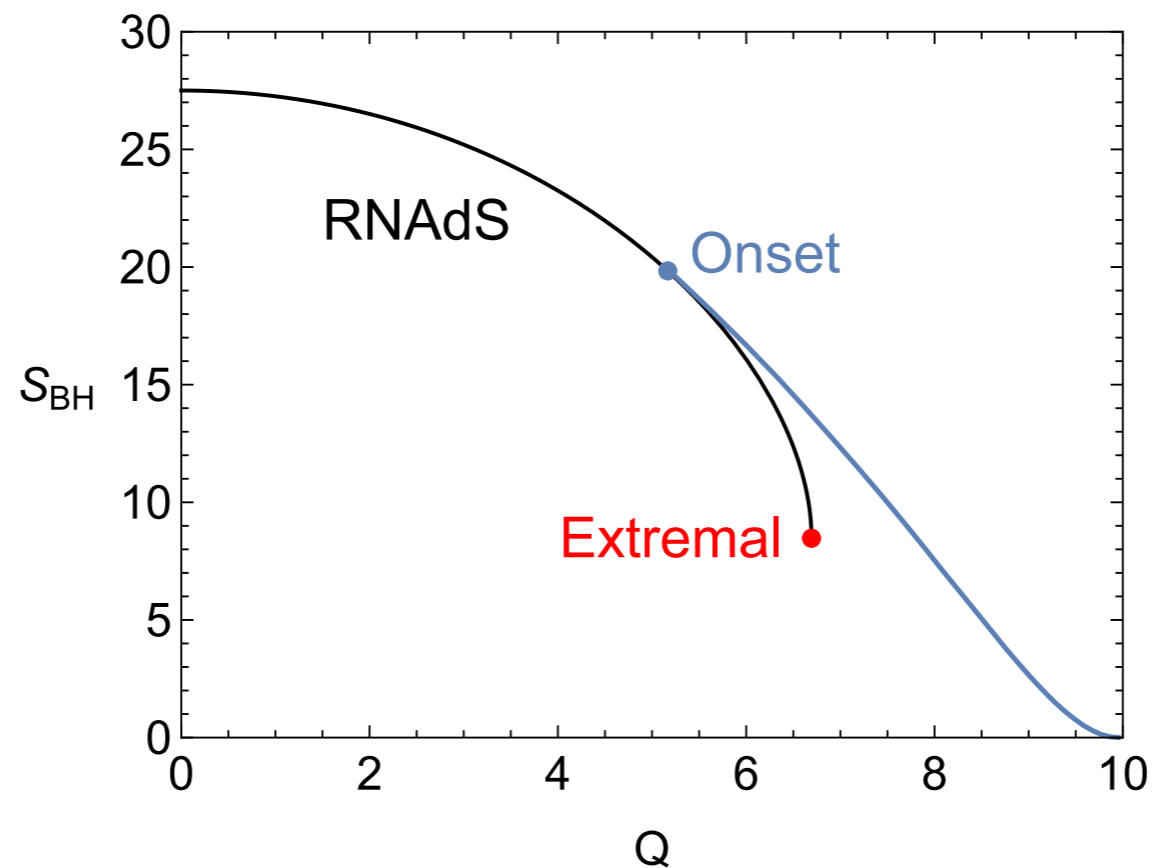


- **When the gauge coupling is small as  $q=1$ , extremal BH can survive as thermodynamically dominant phase.**
- But these structures are not observed if  $q$  is large ( $q > \sqrt{2}$ ).

# Extra: Microcanonical ensemble

For solutions at the same  $(E, Q)$ , a **hairy BH** has **higher entropy** than a zero-scalar BH.

Microcanonical,  $\zeta/\pi=0.6$ ,  $q=1$



# Summary

We constructed hairy BH in global AdS when we use **Robin boundary conditions.**

We studied the **phase diagram in the  $(\mu, T_H)$  plane** when  $(\zeta, q)$  is varied.

Future direction How much can we utilize Robin (mixed) boundary conditions?